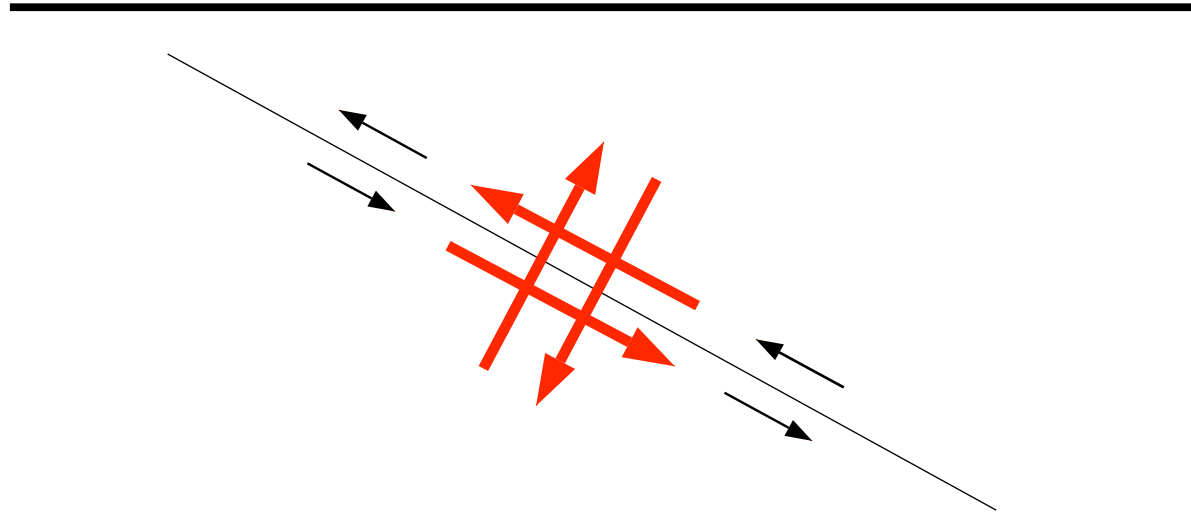


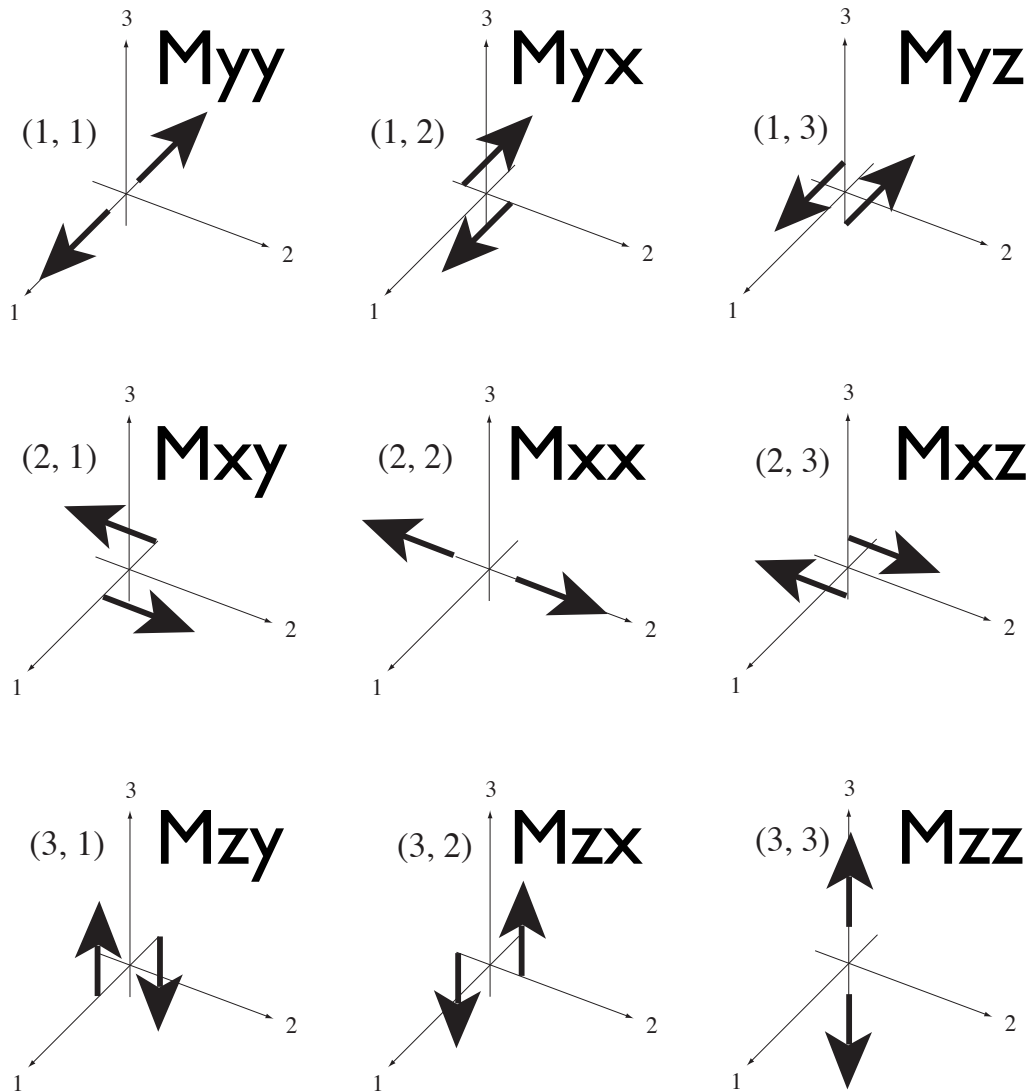
1. Moment-tensor analysis using global data
2. The Global CMT catalog
3. Using calibration information in waveform analysis
4. Data quality control using signals
5. Data quality control using noise
6. Finding interesting things in the noise
7. Using noise for tomography

# Faulting **force** model



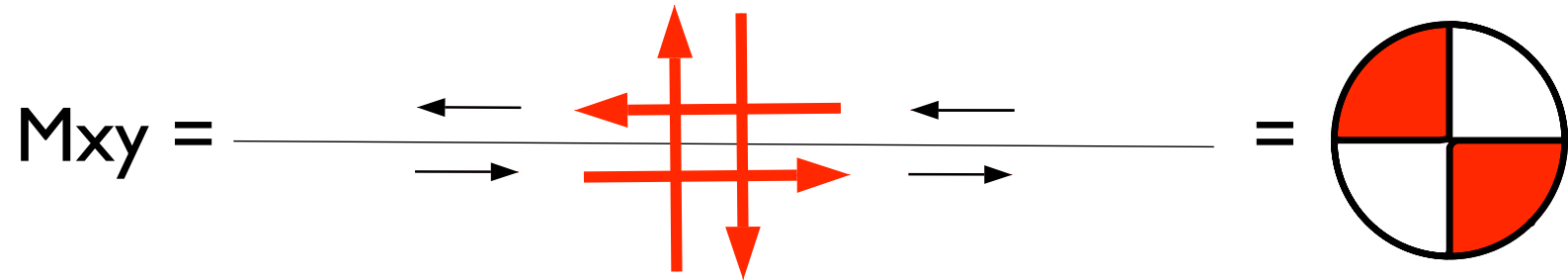
The elastic stress release in an earthquake is described by a double couple of forces

# The nine dipoles of the seismic moment tensor



(Aki and Richards, 2002)

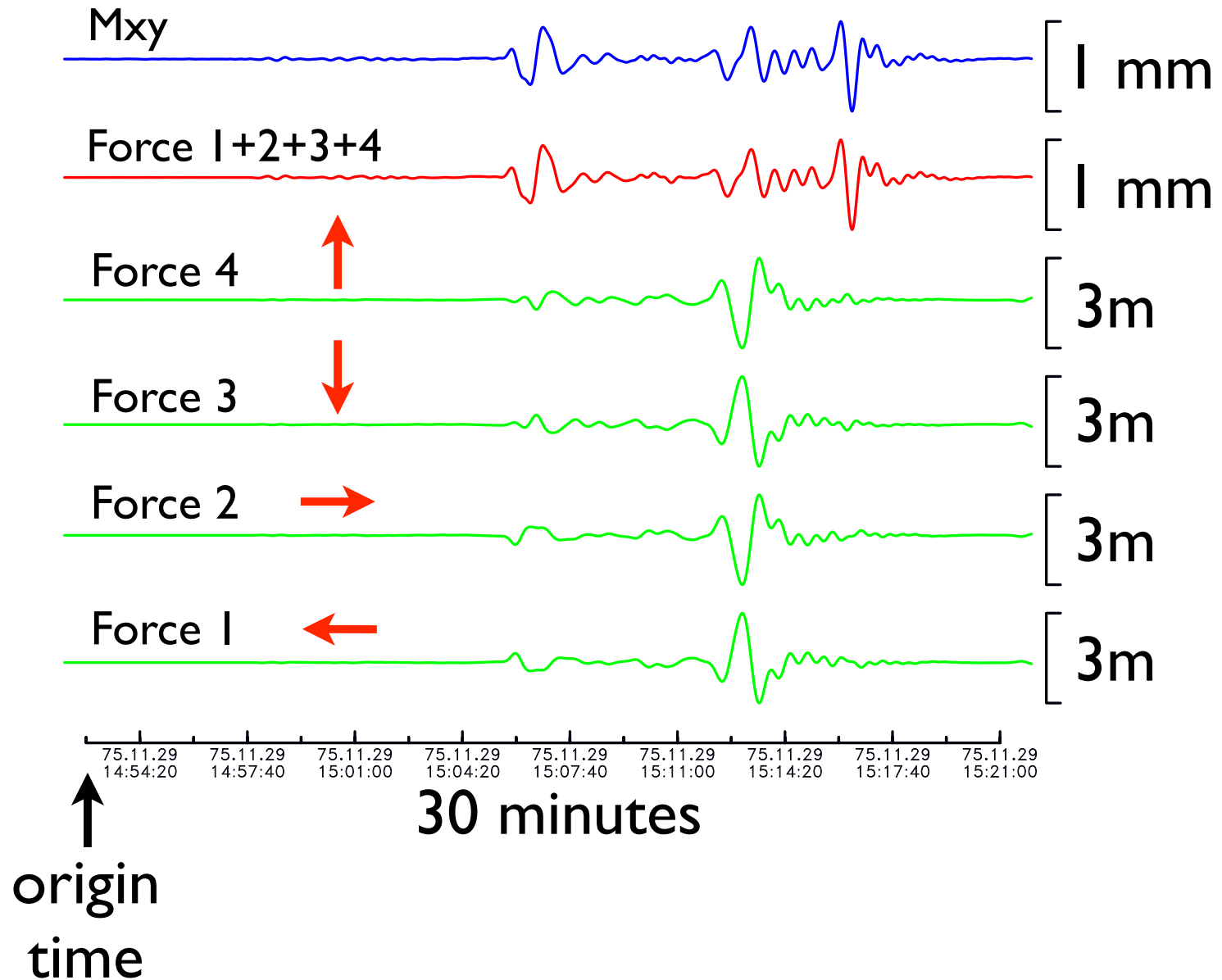
But,  $M_{xy}=M_{yx}$ ,  $M_{yz}=M_{zy}$ ,  $M_{xz}=M_{zx}$



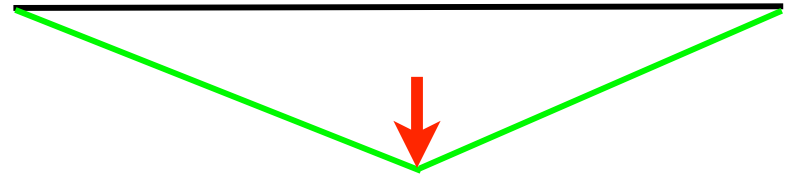
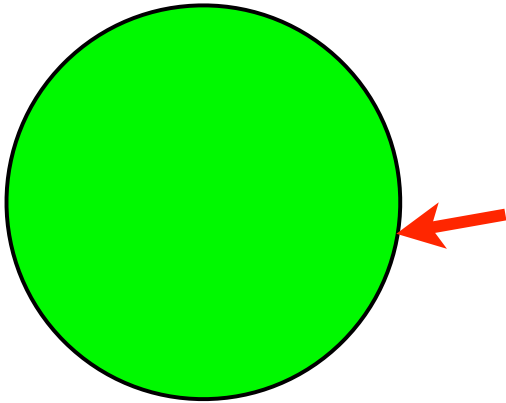
for example,

$$10^{28} \text{ dyne-cm} = 10^{24} \text{ dyne} \times 10000 \text{ cm}$$

# Calculated force seismograms (6000 km distance)



The vibrations caused by a force acting on or in the Earth can be modeled by summation of Earth's normal modes



$$u(\mathbf{x}, t) = \sum_k [1 - \exp[-\alpha_k(t - t_s)] \cos \omega_k(t - t_s)] \mathbf{f} \cdot \mathbf{w}^{(k)}(\mathbf{x}_s) \mathbf{s}_k(\mathbf{x})$$

where  $\mathbf{f}$  is the force vector and  $\mathbf{w}^k$  is the displacement of the  $k$ -th mode.

# Moment-tensor analysis by waveform fitting

(Observed seismogram)/(Instrument response) x Filter = Observed waveform

(Synthetic displacement seismogram) x Filter = Model waveform

Model waveform depends on:

1. Earthquake parameters
2. Earth structure

If the Earth structure and the earthquake location are known, the

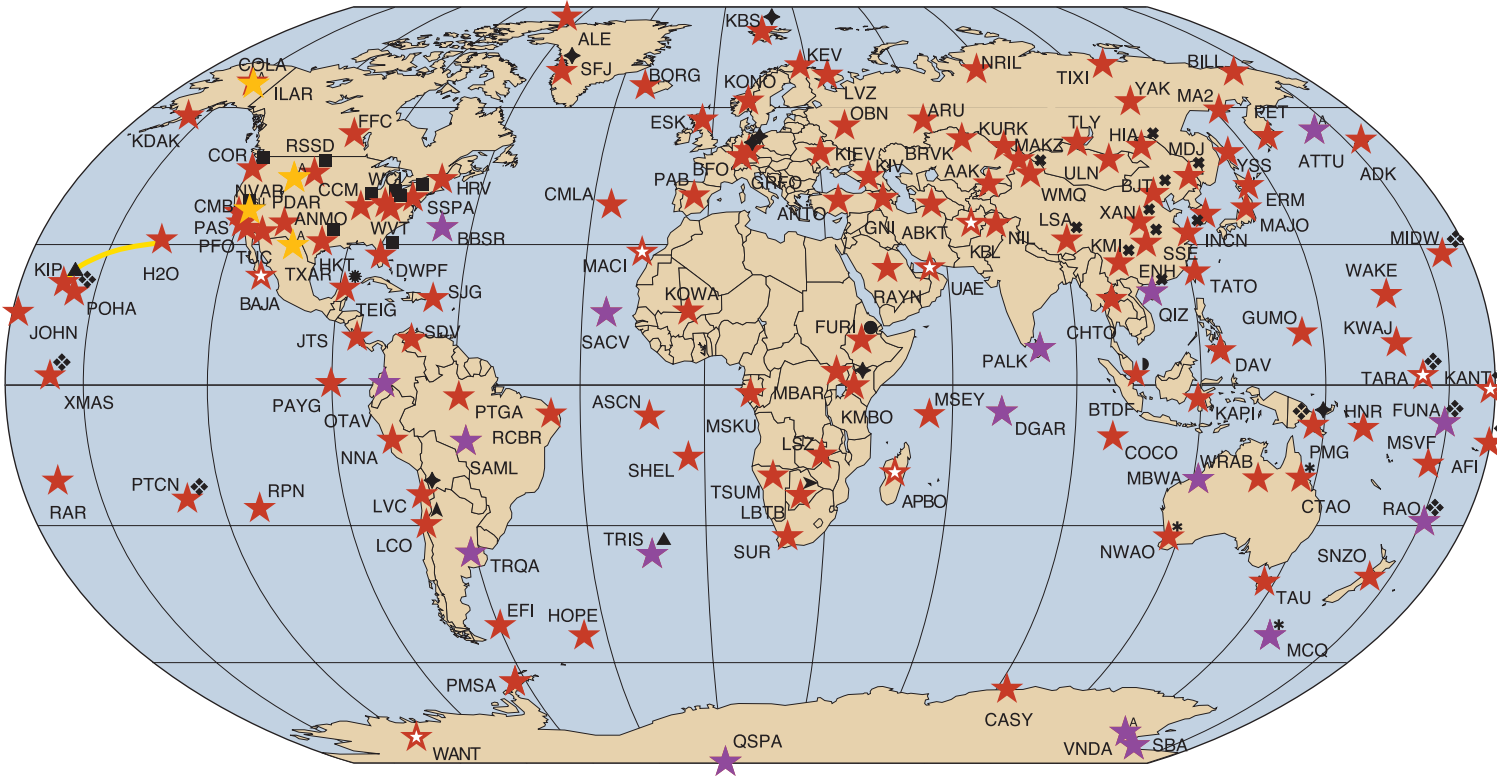
Model waveform depends only on the six elements of the moment tensor,

$M_{xx}$ ,  $M_{yy}$ ,  $M_{zz}$ ,  $M_{xy}$ ,  $M_{xz}$ , and  $M_{yz}$

Minimize the difference  $[\text{Observed waveform} - \text{Model waveform}]^2$

with respect to the moment tensor elements.

# GLOBAL SEISMOGRAPHIC NETWORK



## IRIS International & National Cooperative Sites

IRIS	Affiliate	IRIS International & National Cooperative Sites													
Current	Array	Geoscope	Japan	Mednet	Geofon/AWI/BGR/BFO	China/USGS	Mexico	Singapore	Botswana	Andes	Australia	ANSS	AFTAC	SMU	
★	★	★	★	★	★	★	★	★	★	★	★	★	★	★	

Current GSN station coverage of Earth is shown as of August 2005. Sites added in the past five years are noted in purple (stations) and orange (arrays). Sites planned to be completed are noted with white stars. Cooperative sites are indicated by symbols on the upper right "shoulder" of the stars.



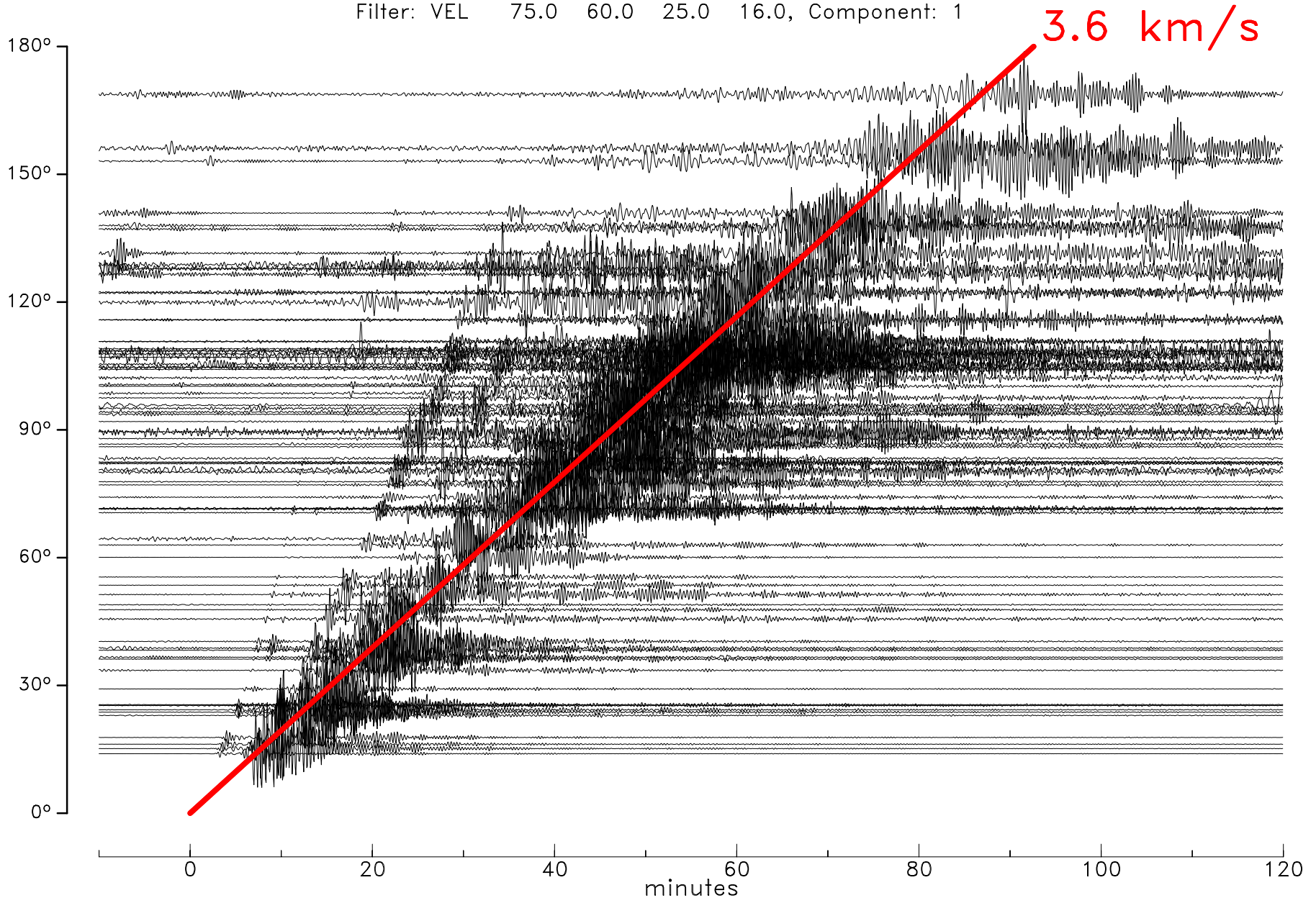
STS-I Seismometer  
at Harvard, Mass.



# Global network record section for an earthquake off the coast of Jalisco, Mexico

E200604040230A

Event: 2006/04/04, 02:30:28.0, OFF COAST OF JALISCO, MEXICO  
Hypocenter (PDE ): Lat= 18.69, Lon=-107.06, h= 33.9, mb=5.9, MS=5.9  
Filter: VEL 75.0 60.0 25.0 16.0, Component: 1



## Moment-tensor analysis by waveform fitting

(Observed seismogram)/(Instrument response) x Filter = Observed waveform

(Synthetic displacement seismogram) x Filter = Model waveform

Model waveform depends on:

1. Earthquake parameters
2. Earth structure

If the Earth structure and the earthquake location are known, the

Model waveform depends only on the six elements of the moment tensor,

$M_{xx}$ ,  $M_{yy}$ ,  $M_{zz}$ ,  $M_{xy}$ ,  $M_{xz}$ , and  $M_{yz}$

Minimize the difference  $[\text{Observed waveform} - \text{Model waveform}]^2$

with respect to the moment tensor elements.

## Seismogram Modeling

The  $k$ -th seismogram in a data set for a given earthquake can be represented by:

$$u_k(\mathbf{r}, t) = \sum_{i=1}^N \psi_{ik}(\mathbf{r}_0, \mathbf{r}, t) f_i$$

where  $\psi_{ik}$  are the excitation kernels and  $f_i$  are independent parameters of the source model.

$f_1 = M_{zz}$ ,  $f_2 = M_{yy}$ , etc.;  $N=6$

## Seismogram Synthesis for a Moment-Tensor Source

The seismic displacement field can be calculated by superposition of the normal modes of the Earth (Gilbert, 1971):

$$u(\mathbf{x}, t) = \sum_k [1 - \exp[-\alpha_k(t - t_s)] \cos \omega_k(t - t_s)] \mathbf{M} : \mathbf{e}^{(k)}(\mathbf{x}_s) \mathbf{s}_k(\mathbf{x})$$

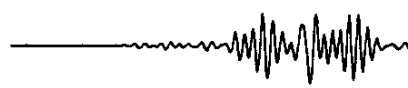
where  $\alpha_k$  is the decay constant of and  $\mathbf{e}^k$  is the strain tensor in the  $k$ -th mode;  $\mathbf{s}_k$  is the eigenfunction of the  $k$ -th mode; and  $\mathbf{M}$  is the seismic moment tensor.

Vertical

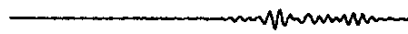
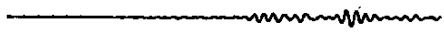
Longitudinal

Transverse

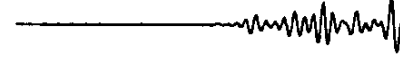
$M_{zz}$   $\psi_1$



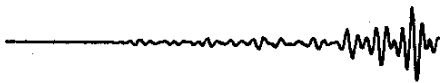
$M_{yy}$   $\psi_2$



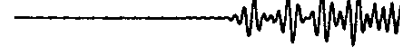
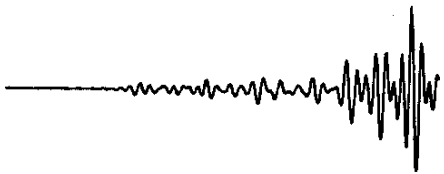
$M_{xx}$   $\psi_3$



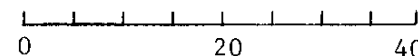
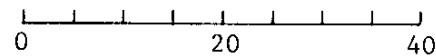
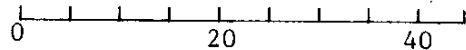
$M_{yz}$   $\psi_4$



$M_{xz}$   $\psi_5$



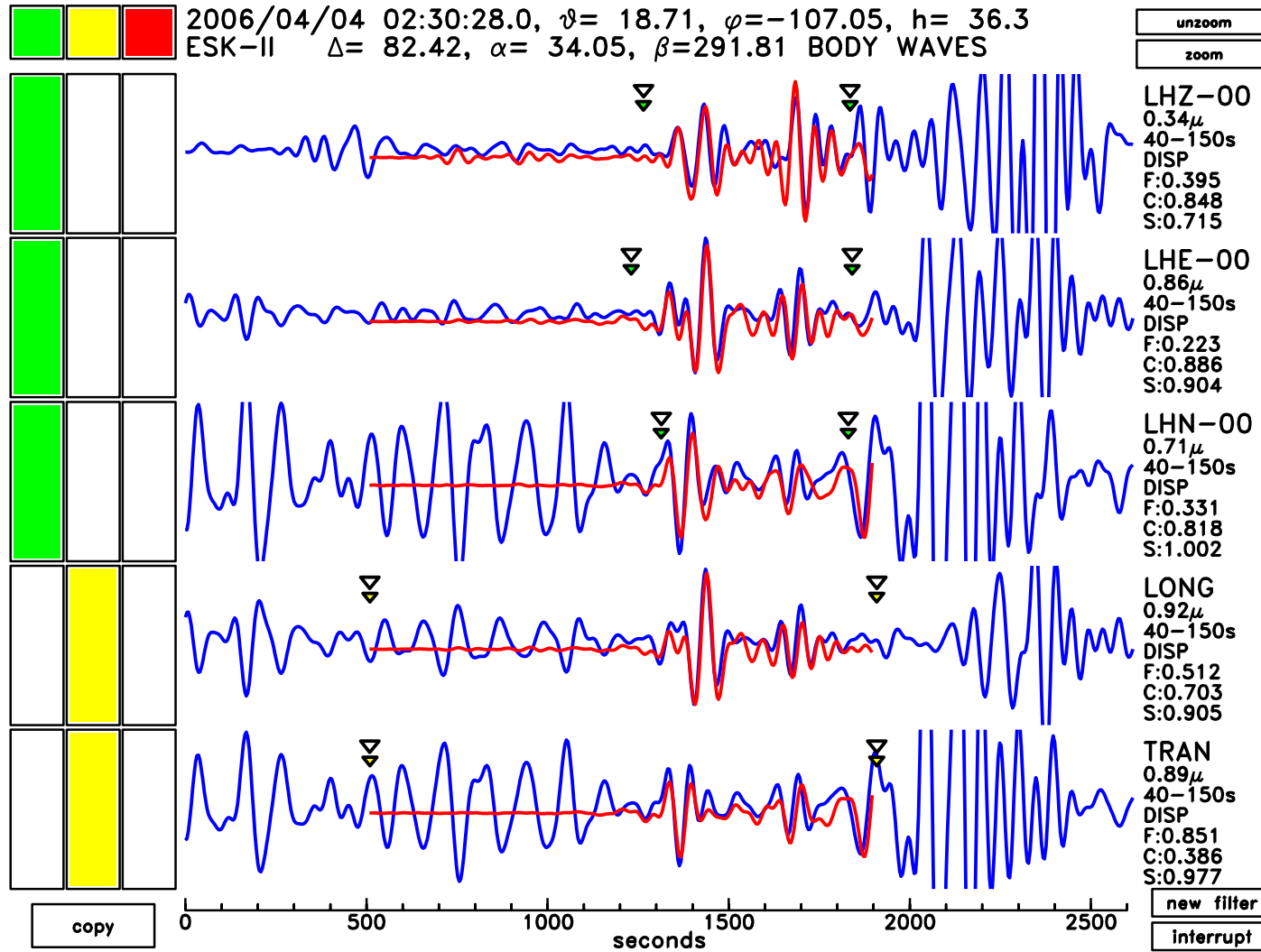
$M_{xy}$   $\psi_6$



Time (minutes)

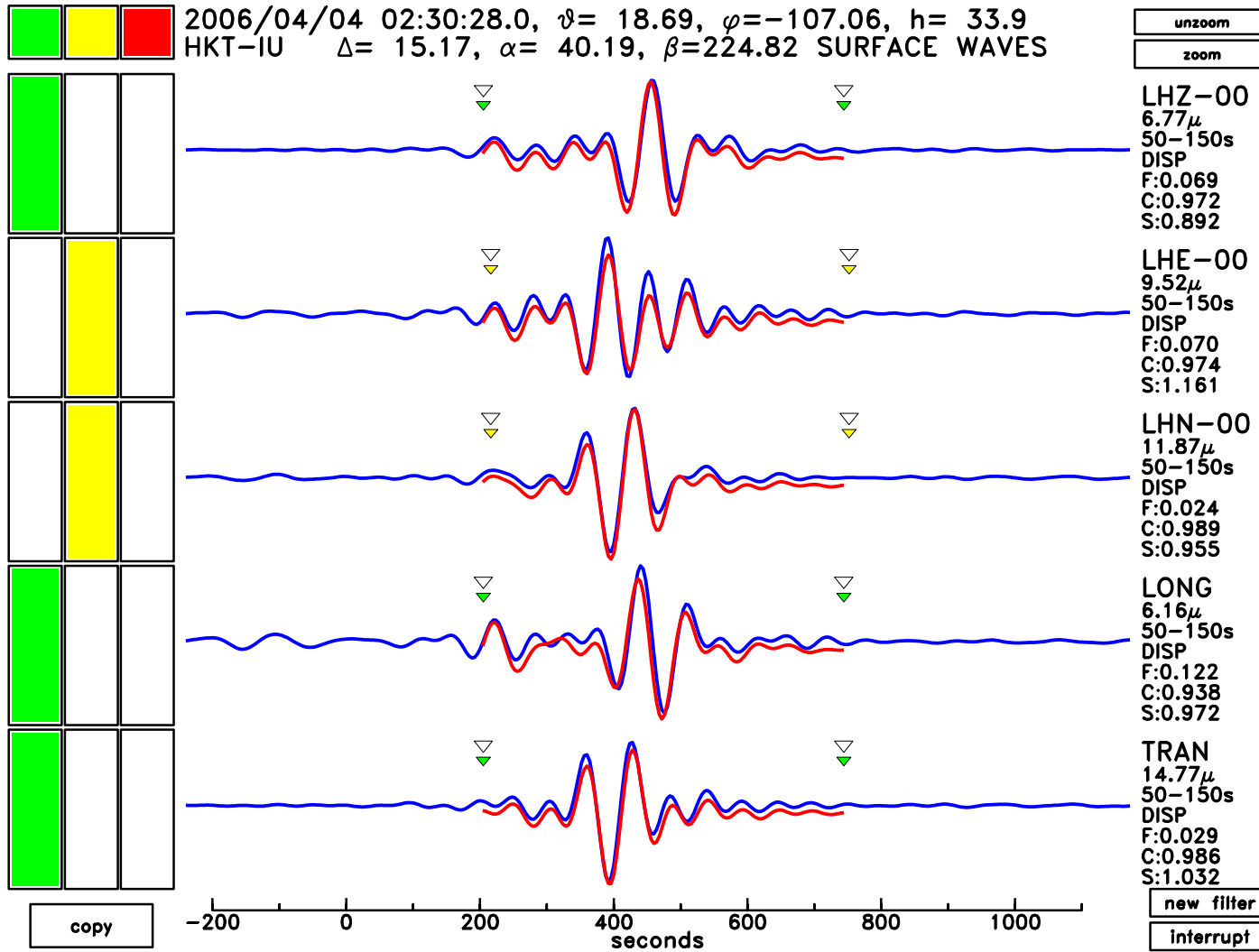
***Excitation kernels for deep earthquake (580 km)***

# Fit to seismograms: Body waves at Eskdalemuir, Scotland



blue - data ; red - model

# Fit to seismograms: Surface waves at Hockley, Texas



blue - data ; red - model



## Estimation of the Source Parameters

For a point source, the elements  $f_i$  can be estimated by solving  $\mathbf{A} \cdot \mathbf{f} = \mathbf{b}$ , where:

$$A_{ij} = \sum_k \int_{t_{k1}}^{t_{k2}} \psi_{ik} \psi_{jk} dt ; b_j = \sum_k \int_{t_{k1}}^{t_{k2}} u_k \psi_{jk} dt.$$

This procedure requires that the position of the source  $(\mathbf{r}_0, t_0)$  be known.

## Solution for the Source Centroid

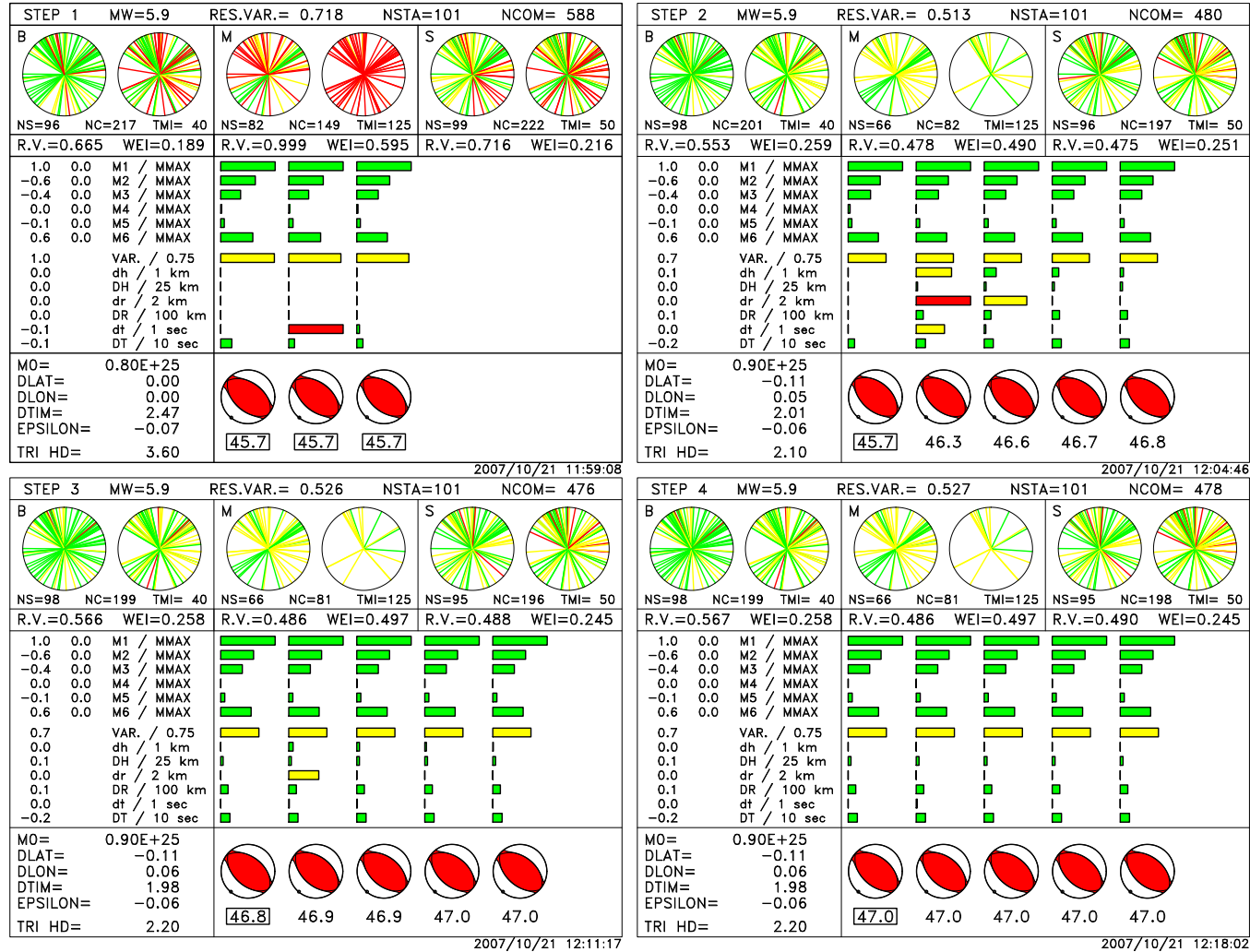
The earthquake centroid can be determined simultaneously with the source model parameters by expansion of the equations of condition to allow for a perturbation in the location of the source (Dziewonski, Chou and Woodhouse, 1981):

$$u_k = u_k^{(0)} + \{\psi_{ki,j}^{(0)} \cdot \delta x_j - \psi_{ki,t}^{(0)} \cdot \delta t_0\} \cdot f_i^{(0)} + \psi_{ki}^{(0)} \cdot \delta f_i ;$$

where the superscript (0) indicates parameters determined for the starting location. The problem can then be solved iteratively.

# Iterative procedure for moment-tensor source converges nicely

Event: 2007/10/21, 10:24:54.0, BOUGAINVILLE REGION, P.N.G.  
 E200710211024A Hypocenter (PDE): Lat= -6.42, Lon= 154.70, h= 45.7, mb=6.2, MS=6.2  
 Centroid : Lat= -6.53, Lon= 154.76, h= 47.0, MW=5.9





2. The Global CMT catalog

3. Using calibration information in waveform analysis

4. Data quality control using signals

5. Data quality control using noise

6. Finding interesting things in the noise

7. Using noise for tomography

## The Global CMT Project

Project started in 1981 (A.M. Dziewonski et al.)

Goal is now to determine source parameters for all earthquakes with  $M > 5$  worldwide

CMT catalog contains ~41,000 moment tensors for the period 1976-2014

In 2006 the project moved from Harvard University to Lamont-Doherty Earth Observatory at Columbia University

The CMT catalog can be accessed at  
[www.globalcmt.org](http://www.globalcmt.org)

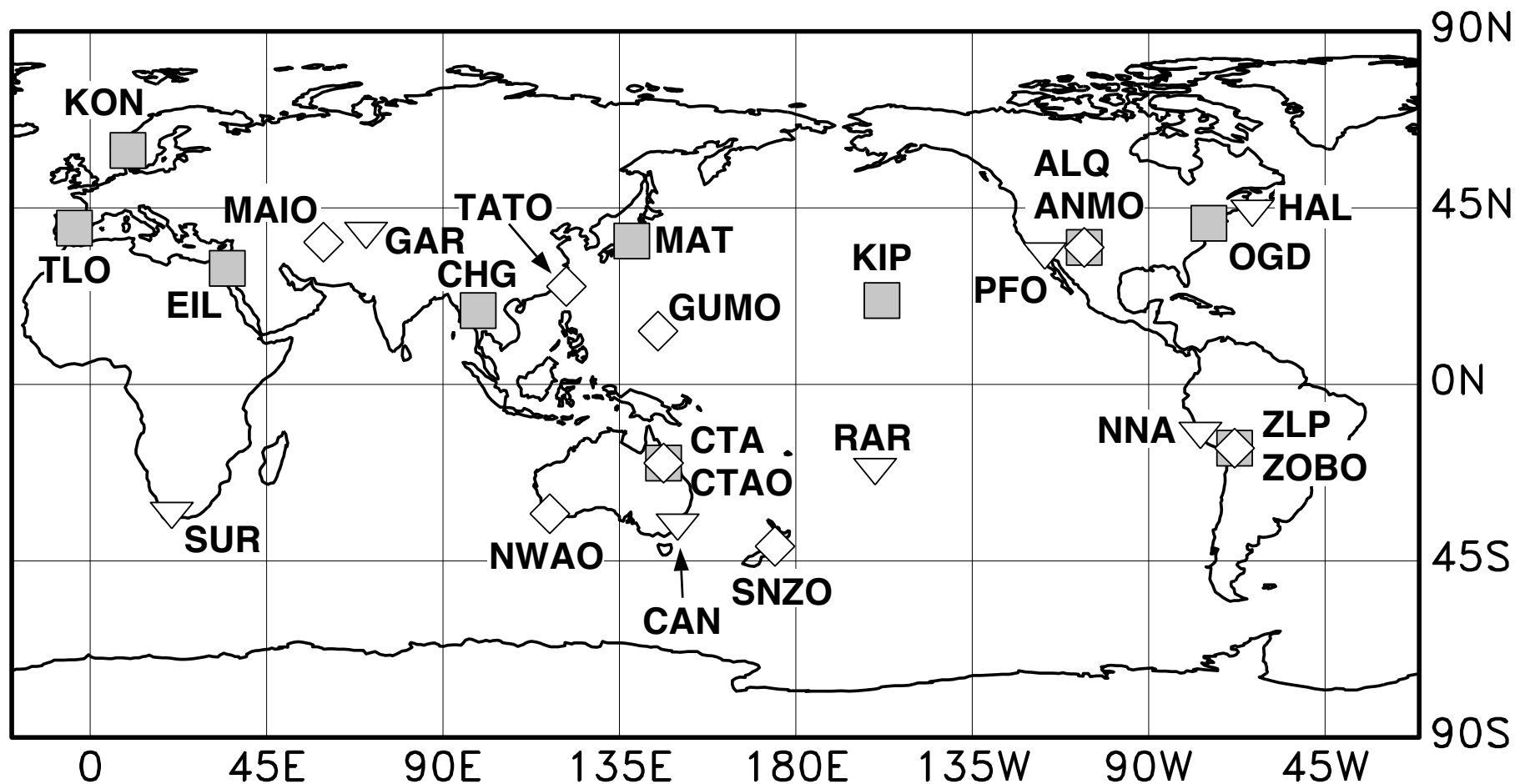
To receive Quick CMT solutions by email,  
send me an email at  
[ekstrom@Ideo.columbia.edu](mailto:ekstrom@Ideo.columbia.edu)

3. Using calibration information in waveform analysis
4. Data quality control using signals
5. Data quality control using noise
6. Finding interesting things in the noise
7. Using noise for tomography



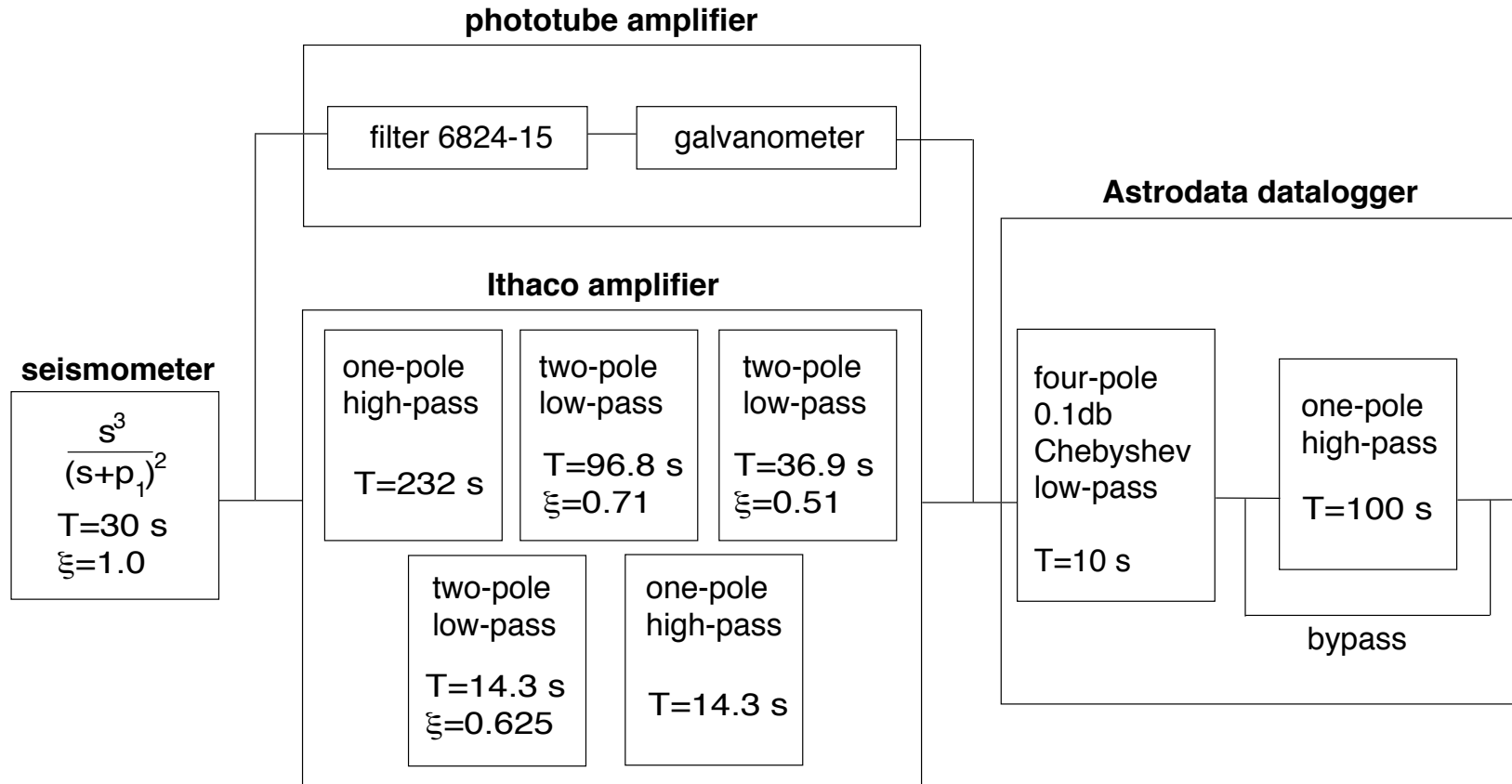
Quantitative waveform analysis requires  
highly accurate instrument response information

# The Global Digital Network in 1976

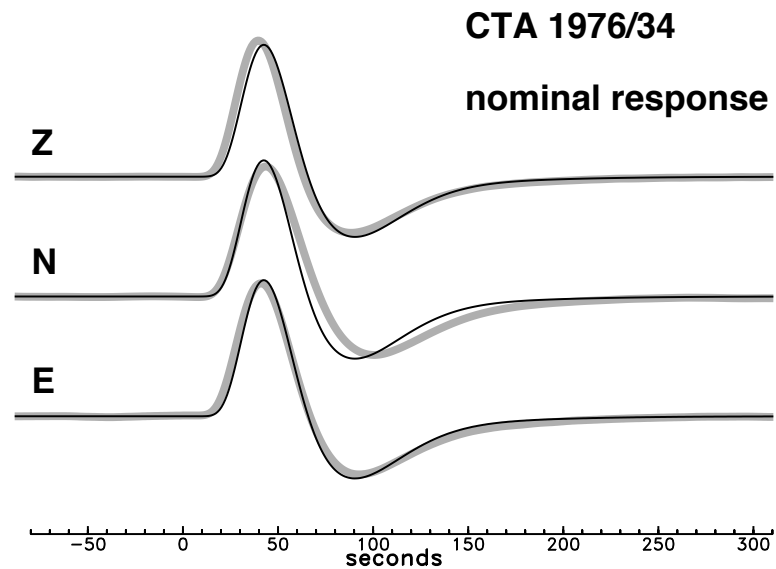


■ High-Gain Long-Period (HGLP) network

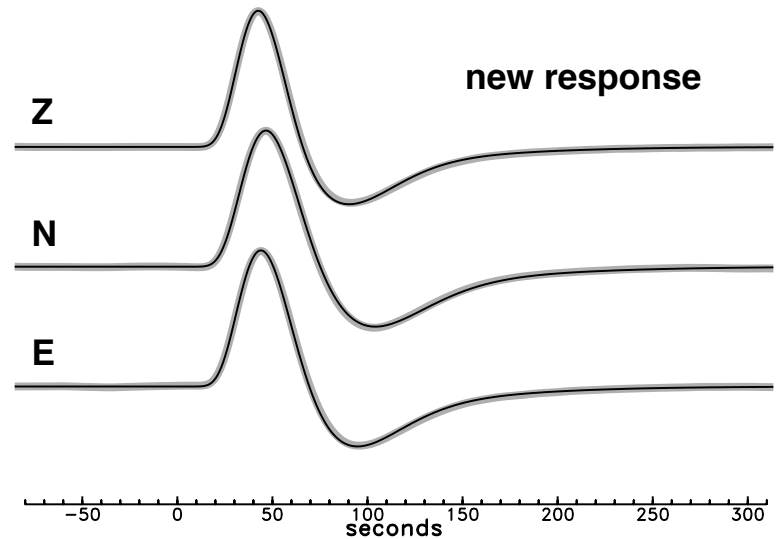
# HGLP seismometer and recording system



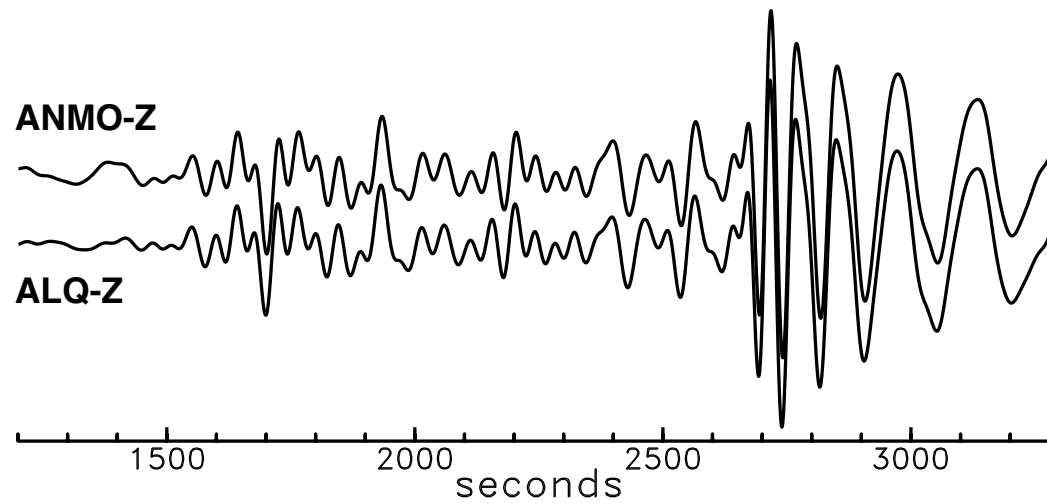
Original calibration  
pulses and pulses for  
nominal response



Original calibration  
pulses and pulses for  
new response after  
inversion

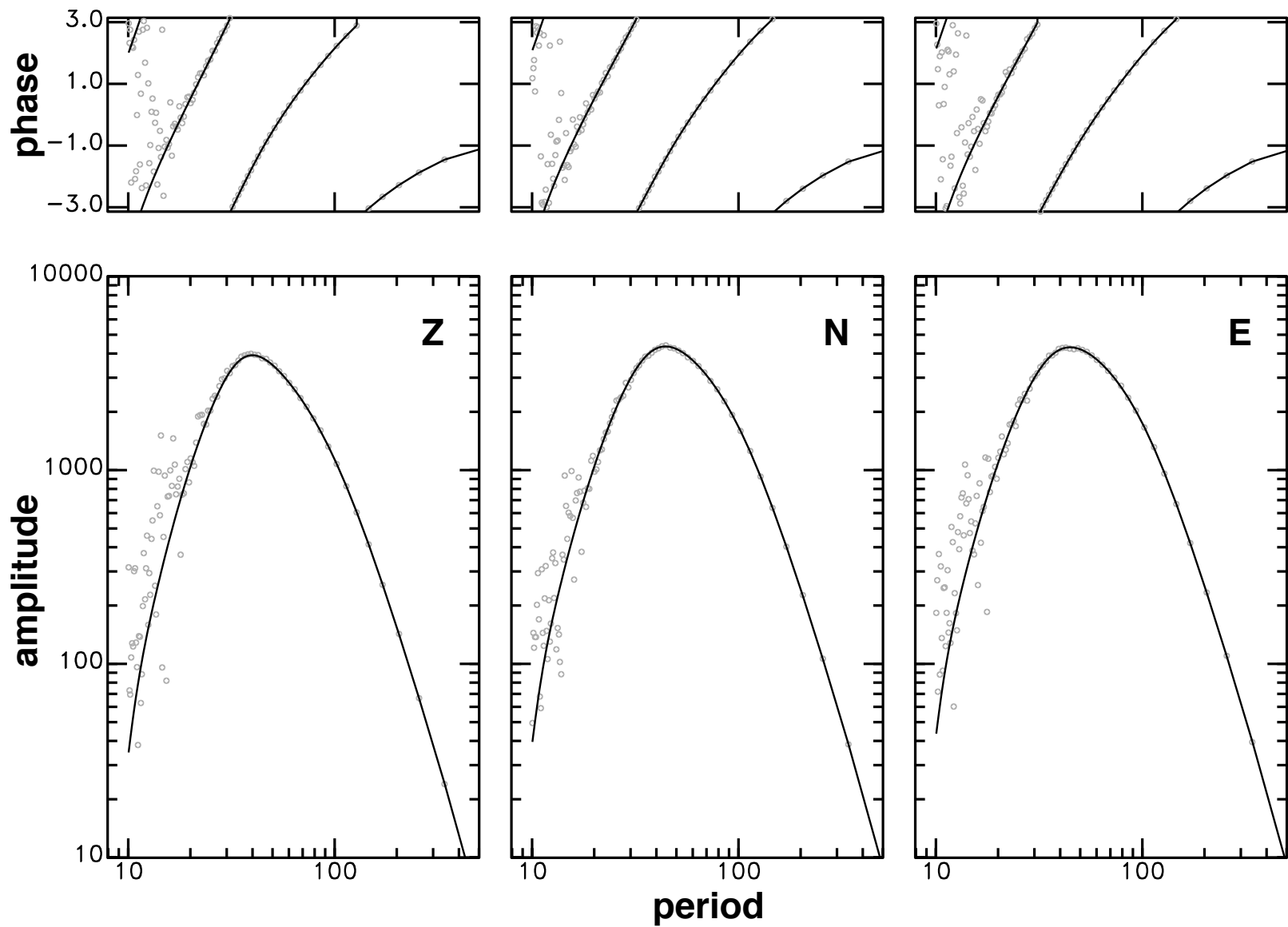


# Comparison of waveforms after normalizing responses for two stations in the same location

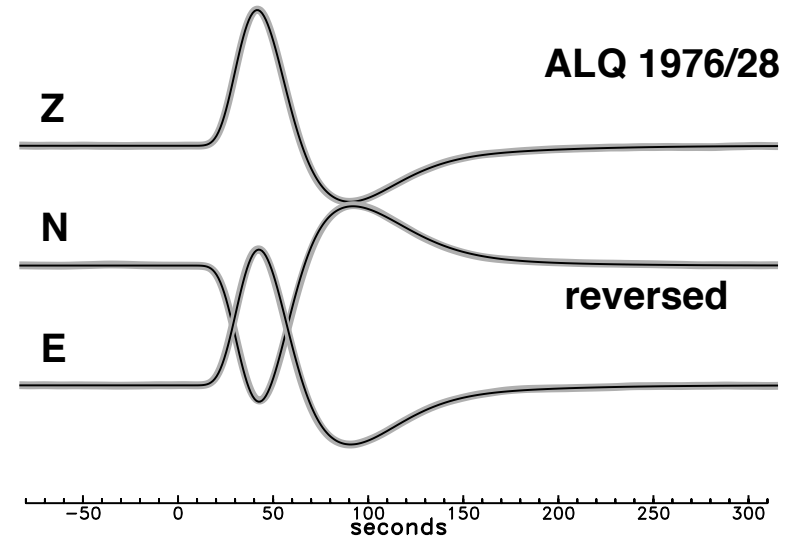


# Check of new responses -- sine-wave calibrations

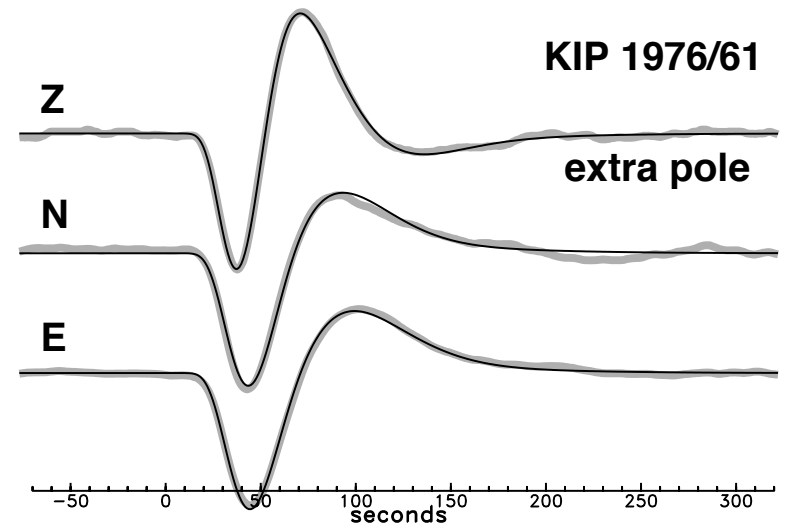
KON 1976/33



Some channels were reversed for some periods of time

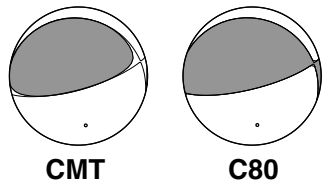


Some channels had extra filters for some periods of time

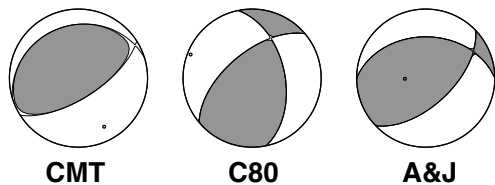


# Waveform comparisons (observed and synthetic) after correcting seismograms using new responses: The 1976 Friuli earthquake

## Friuli Events

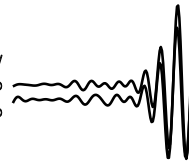


**Main Shock**  
6 May 1976

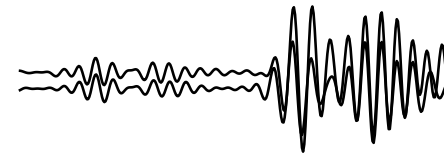


**Aftershock**  
15 Sept. 1976

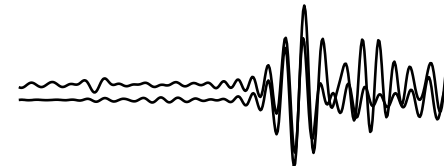
MAIO E-W  
Dist. 35.7°  
Azim. 89.5°



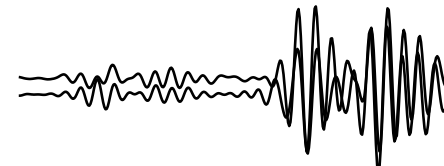
ALQ Vert.  
Dist. 82.6°  
Azim. 314.0°



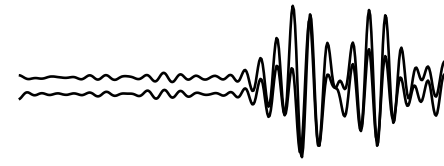
ALQ E-W  
Dist. 82.6°  
Azim. 314.0°



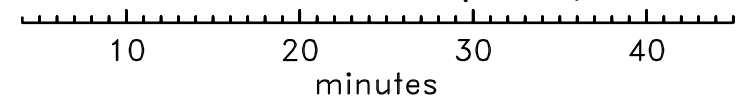
MAT Vert.  
Dist. 83.8°  
Azim. 41.6°



MAT N-S  
Dist. 83.8°  
Azim. 41.6°



KIP N-S  
Dist. 112.1°  
Azim. 351.2°





# Main Point:

Quantitative waveform analysis requires highly accurate instrument response information

4. Data quality control using signals

5. Data quality control using noise

6. Finding interesting things in the noise

7. Using noise for tomography

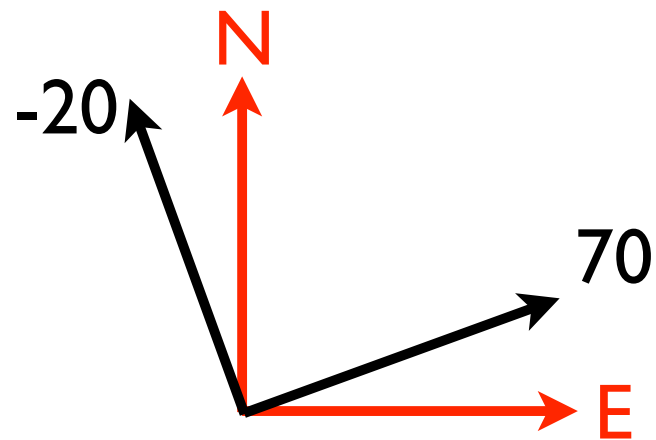
**4a. Sensor orientation**

**4b. Sensor response stability**

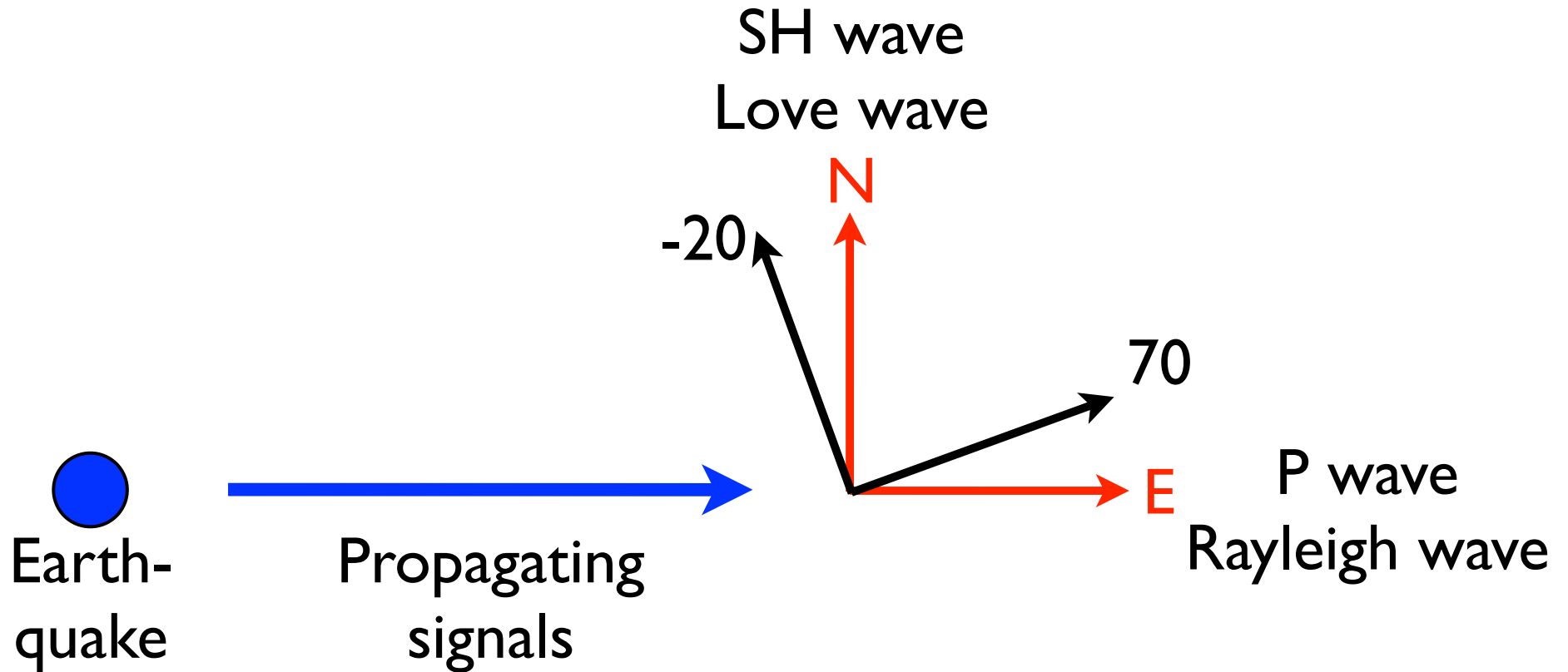
## Horizontal Polarization Problems

Desired (assumed) orientation of seismometer

True orientation of seismometer

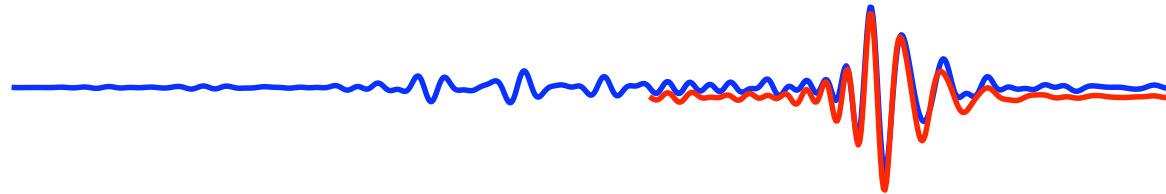


# Natural Polarization of Earthquake Signals



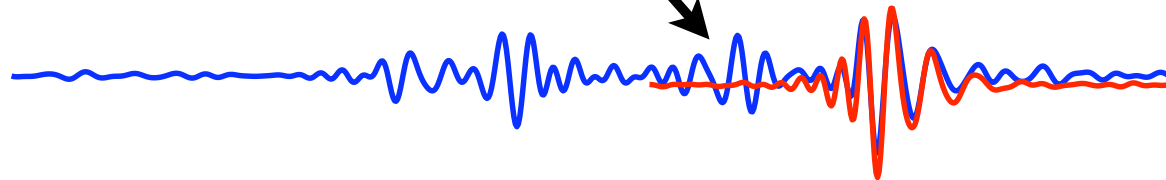
# Symptoms of a misoriented sensor

Vertical



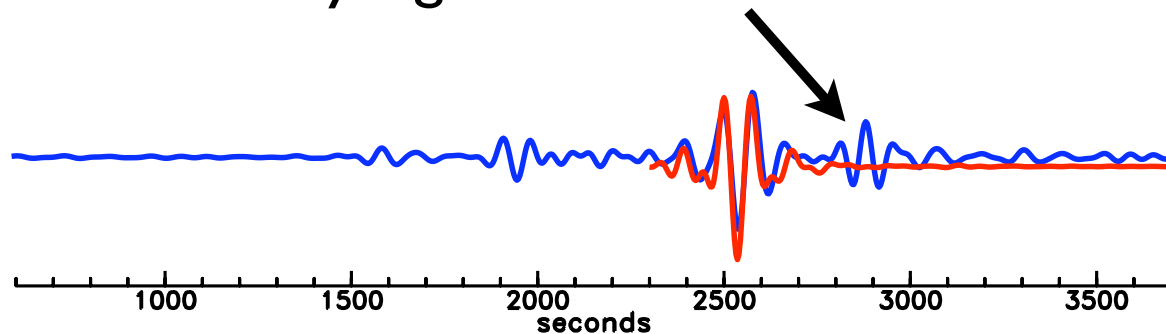
*Love wave on longitudinal*

Longitudinal



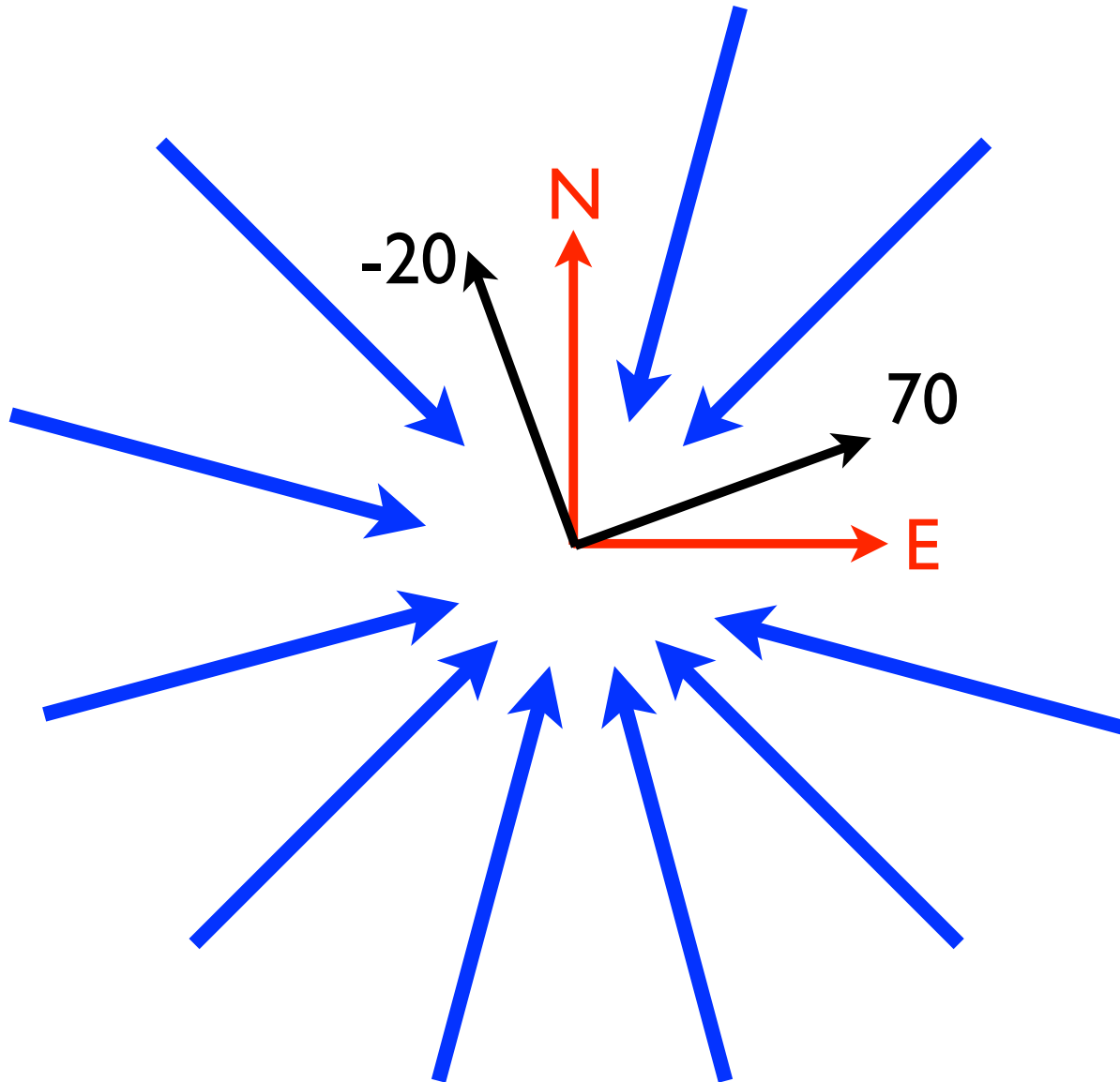
*Rayleigh wave on transverse*

Transverse

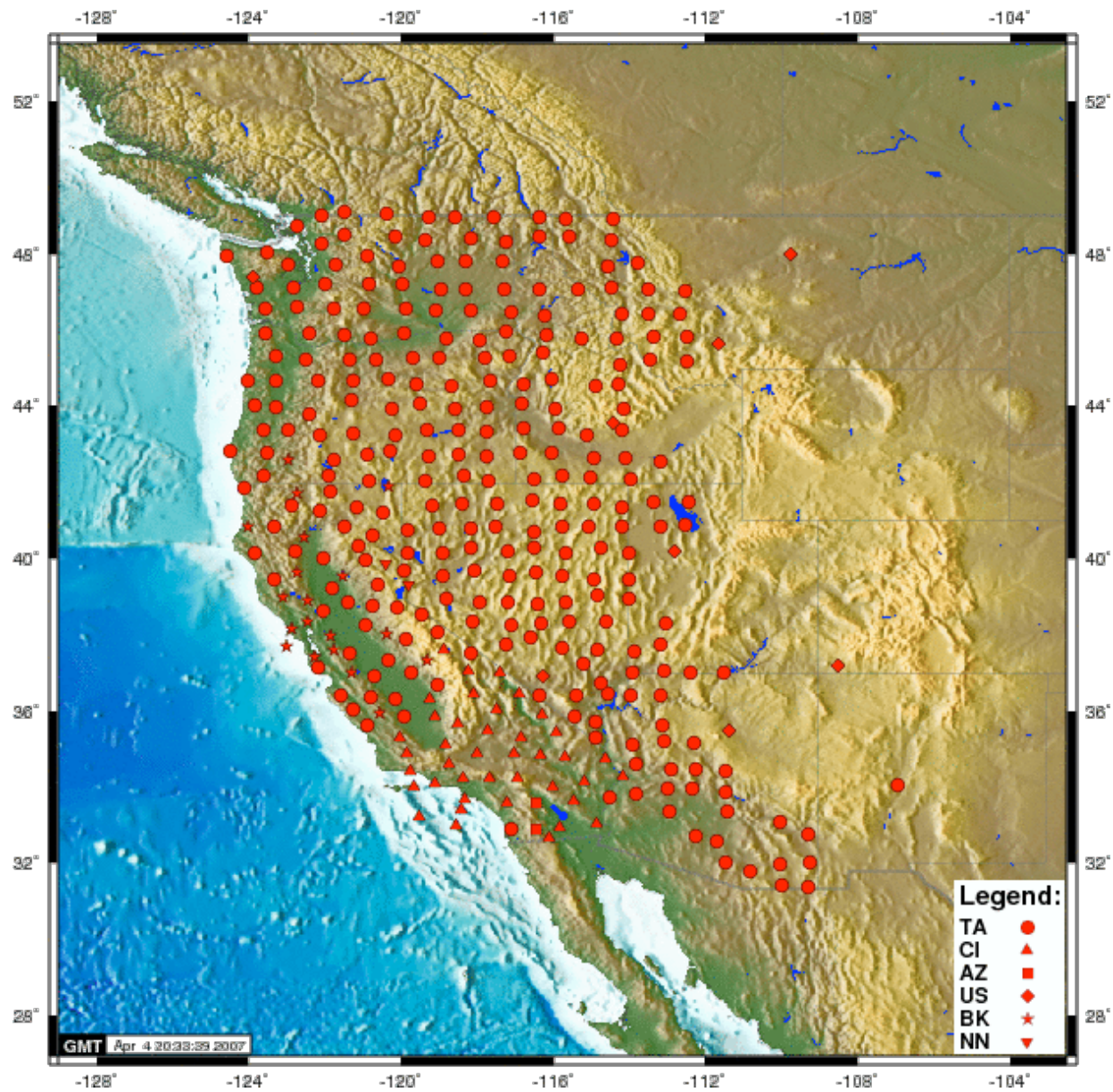


Station D09A, earthquake on 08/20/2007

Many earthquake signals --  
invert for orientation of sensor



# USArray Transportable Array, April 2007



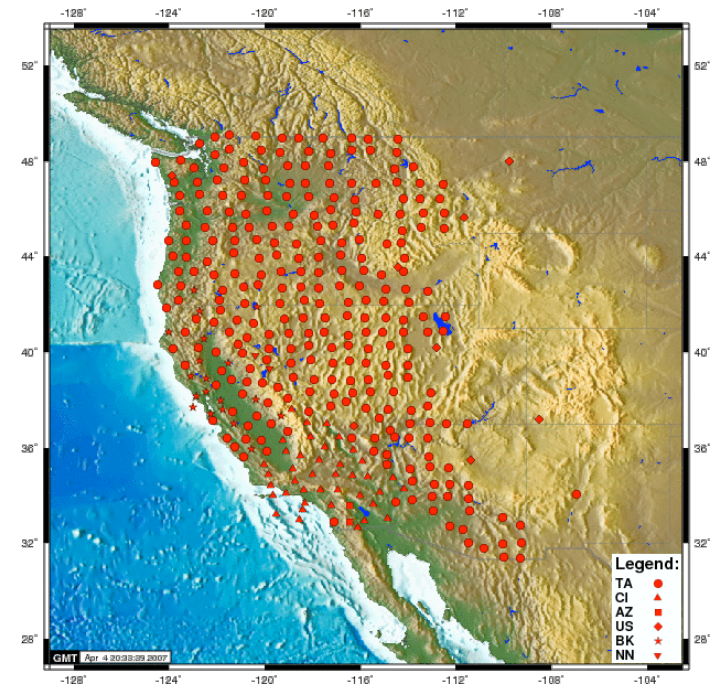


# Polarization analysis of USArray data using earthquake signals

400+ USArray stations

Result:

- > 5% misoriented > 10 degrees
- > 10 % misoriented > 5 degrees



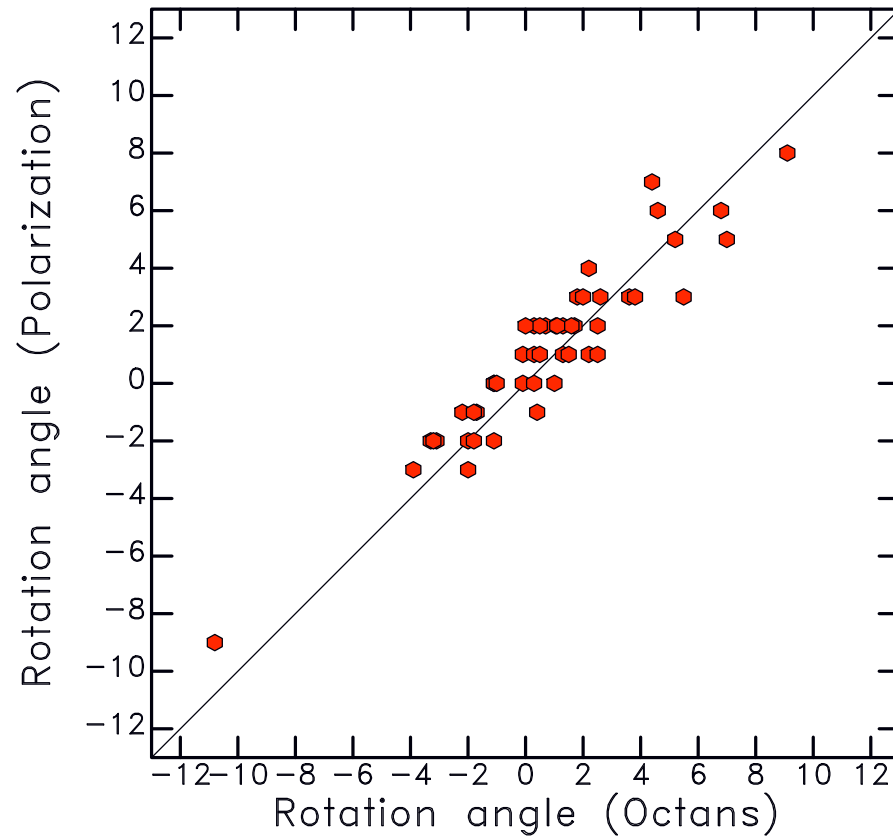
This is a common problem in many networks!

# Octans interferometric laser gyro



# Agreement of field (Octans) and polarization angles

estimated from  
seismograms



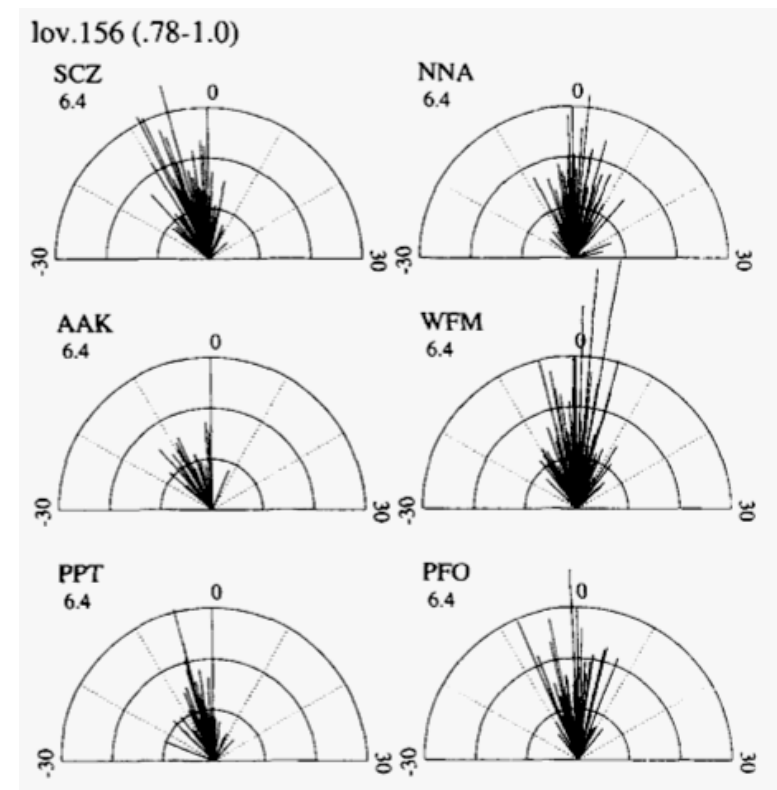
measured in the field

# Sensor orientation

Most GSN and USArray TA stations are well oriented,  
but not all.

## Why does it matter?

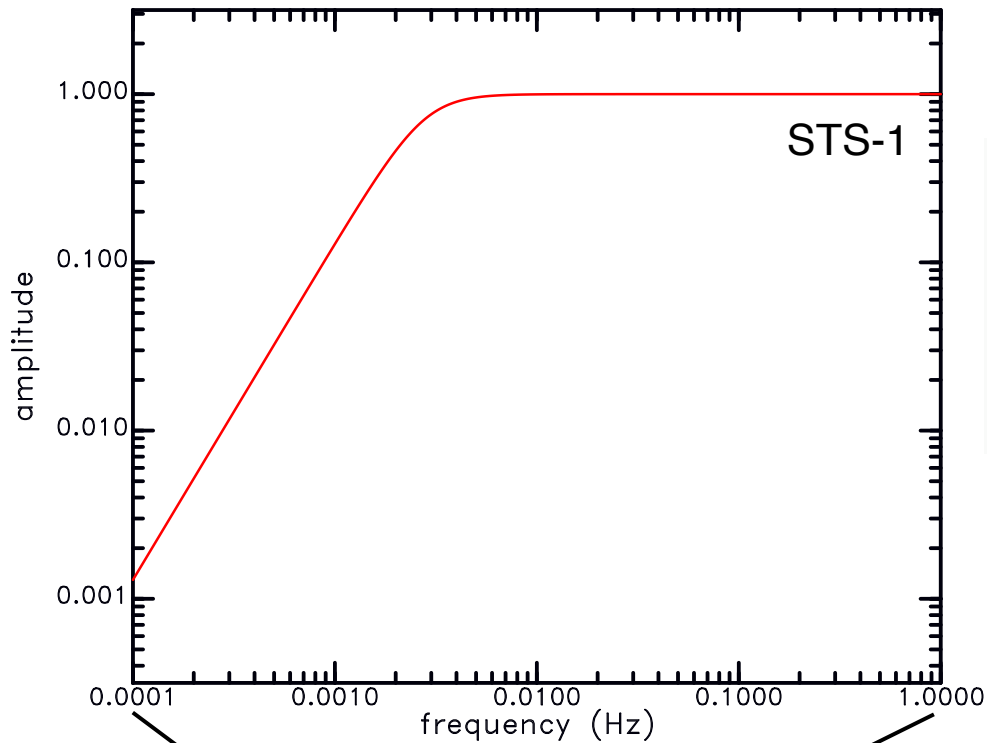
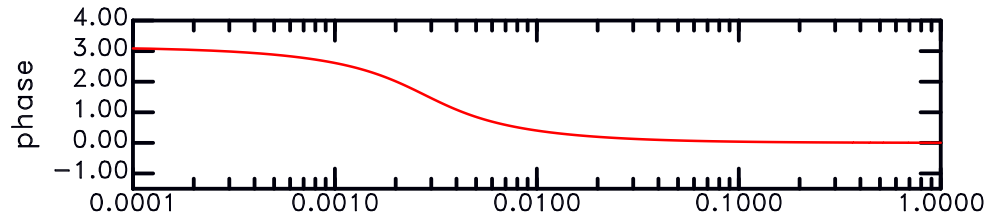
- Modeling of earthquake sources
- Measurement of Love wave / toroidal mode parameters
- Estimates of anisotropy
- Estimates of off-great-circle arrival angle, for both elastic and anelastic structure (tomography)



(Laske, 1995)

## 4b. Sensor response stability

# Seismometer frequency response



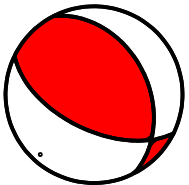
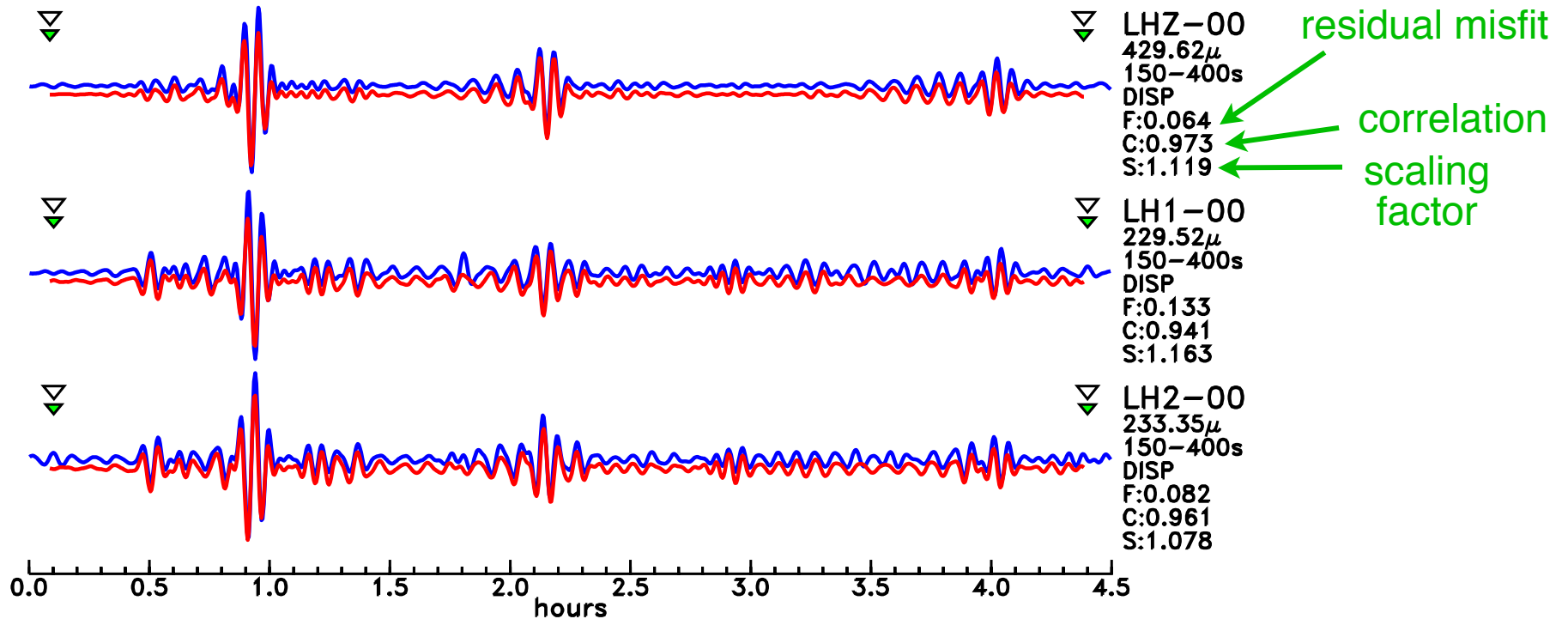
$$T(s) = K \frac{\prod_{i=1}^N (s - z_i)}{\prod_{j=1}^M (s - p_j)}$$



Blue - observed seismograms

Red - synthetic seismograms

2005/10/08 03:50:38.0,  $\vartheta = 34.43$ ,  $\varphi = 73.54$ ,  $h = 10.0$   
POHA-IU  $\Delta = 108.72$ ,  $\alpha = 48.71$ ,  $\beta = 318.75$  MANTLE WAVES

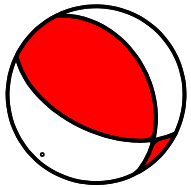
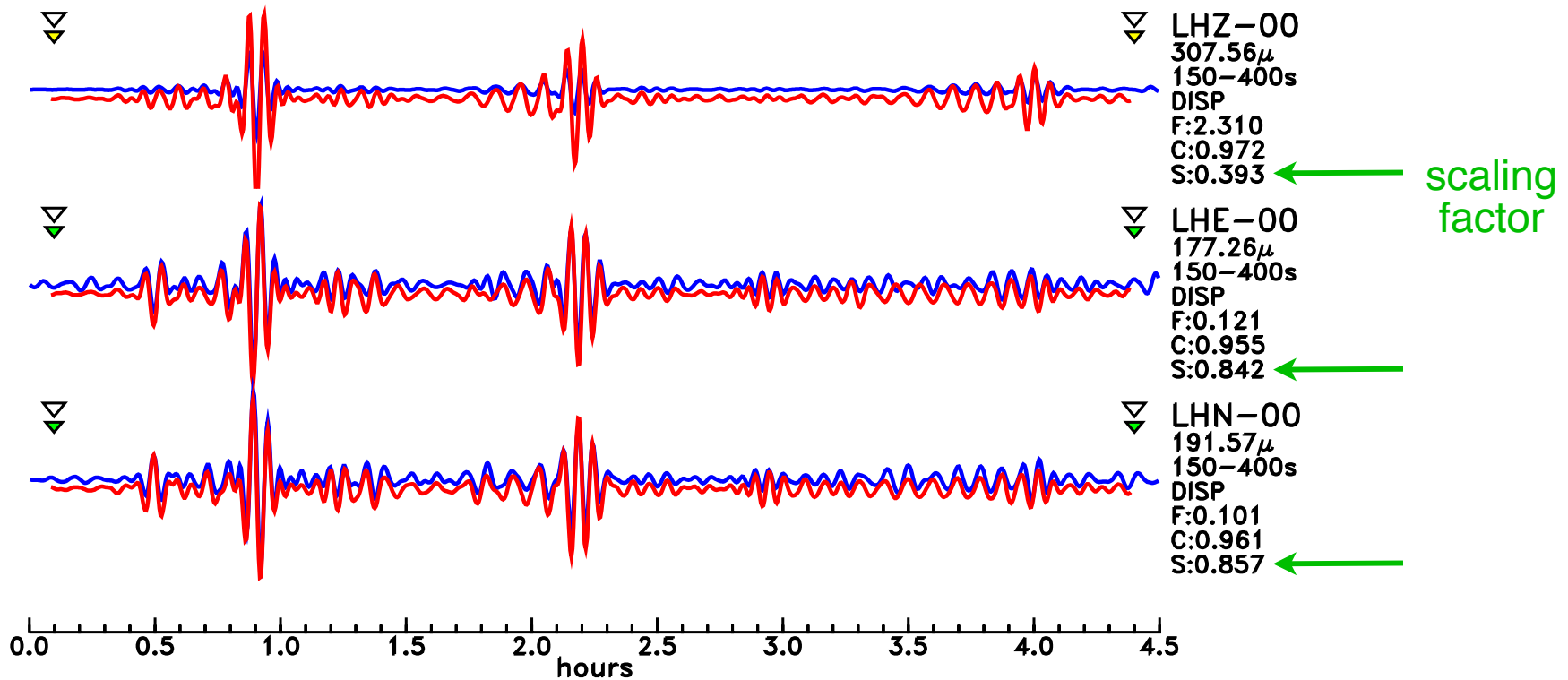


$$S = \frac{\sum_{i=1}^N O_i S_i}{\sum_{i=1}^N S_i^2}$$

Blue - observed seismograms

Red - synthetic seismograms

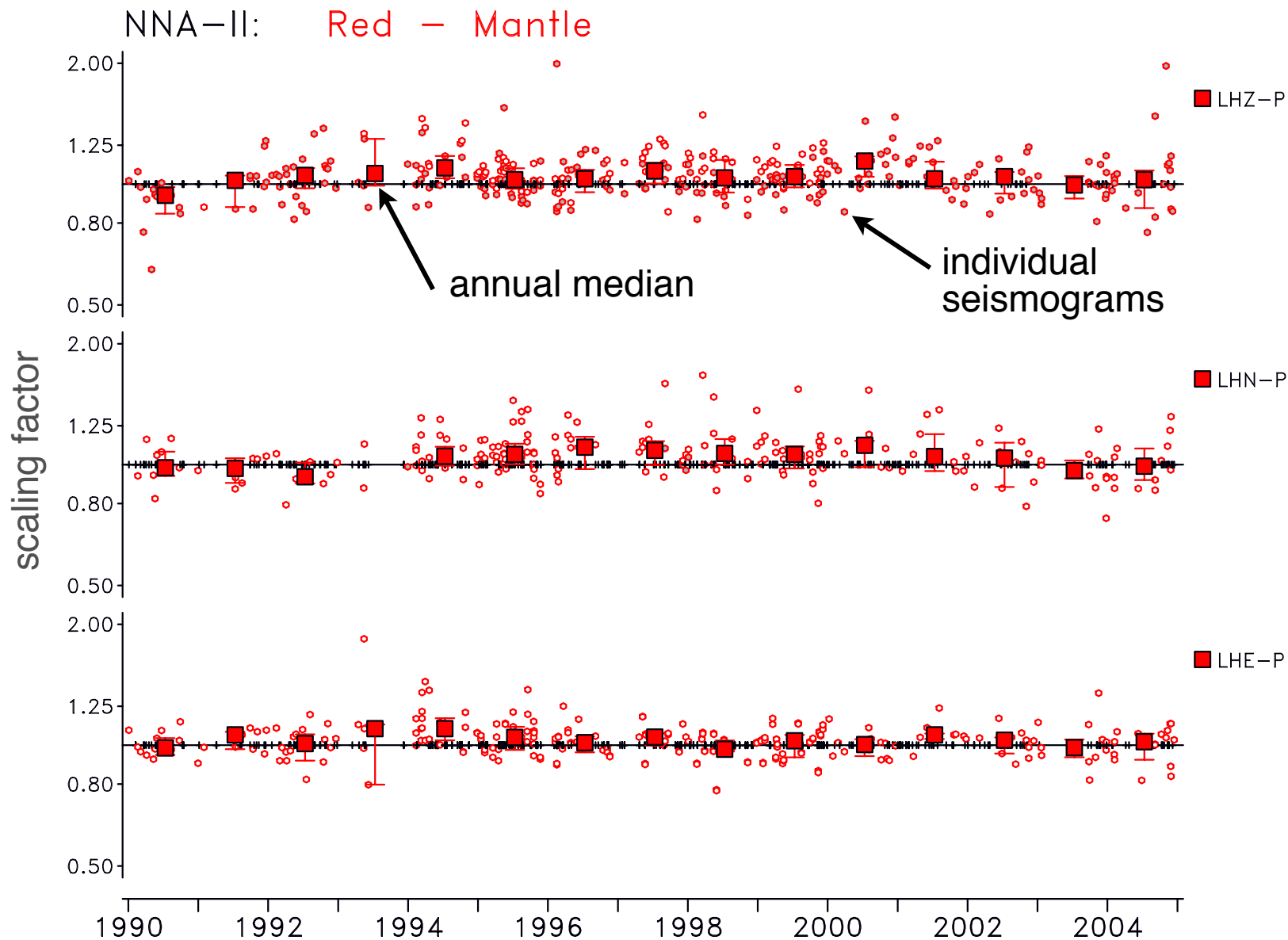
2005/10/08 03:50:38.0,  $\vartheta = 34.43$ ,  $\varphi = 73.54$ ,  $h = 10.0$   
KIP-IU  $\Delta = 105.93$ ,  $\alpha = 49.37$ ,  $\beta = 317.68$  MANTLE WAVES



$$S = \frac{\sum_{i=1}^N O_i S_i}{\sum_{i=1}^N S_i^2}$$

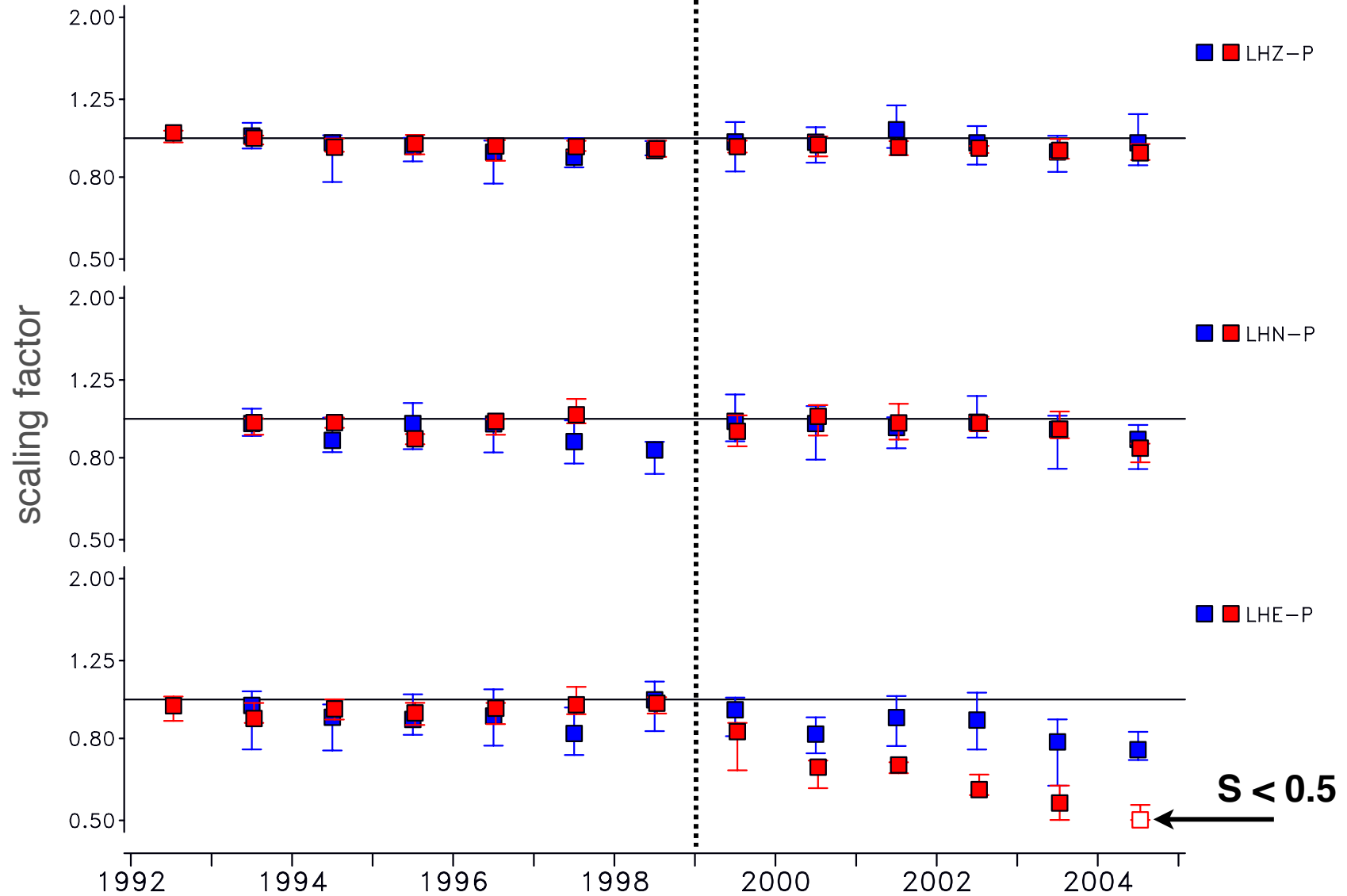


# Scaling factors at NNA-II, 1990-2004



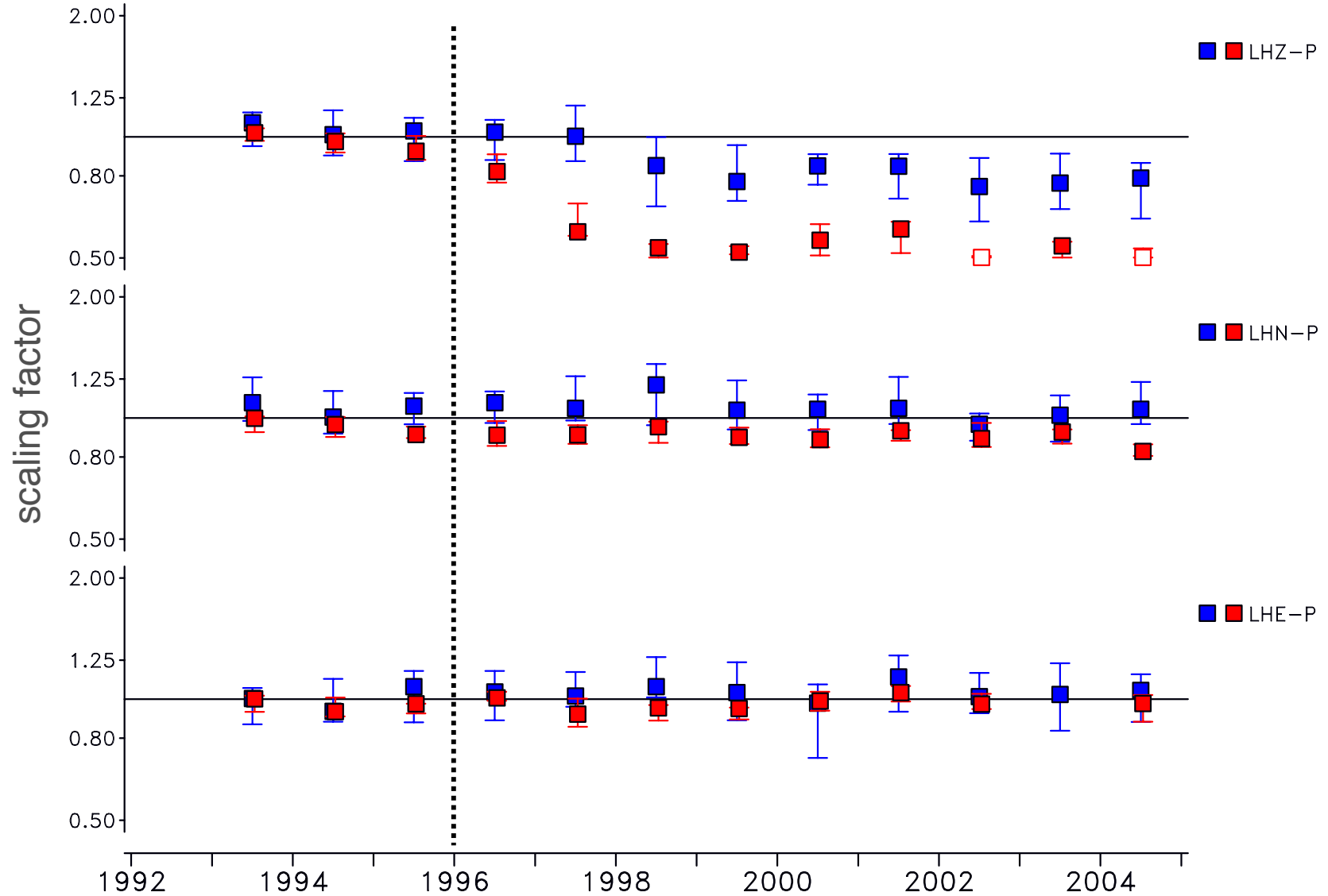
# Scaling factors at PAB-IU, 1992-2004

PAB-IU: Blue — Body; Red — Mantle



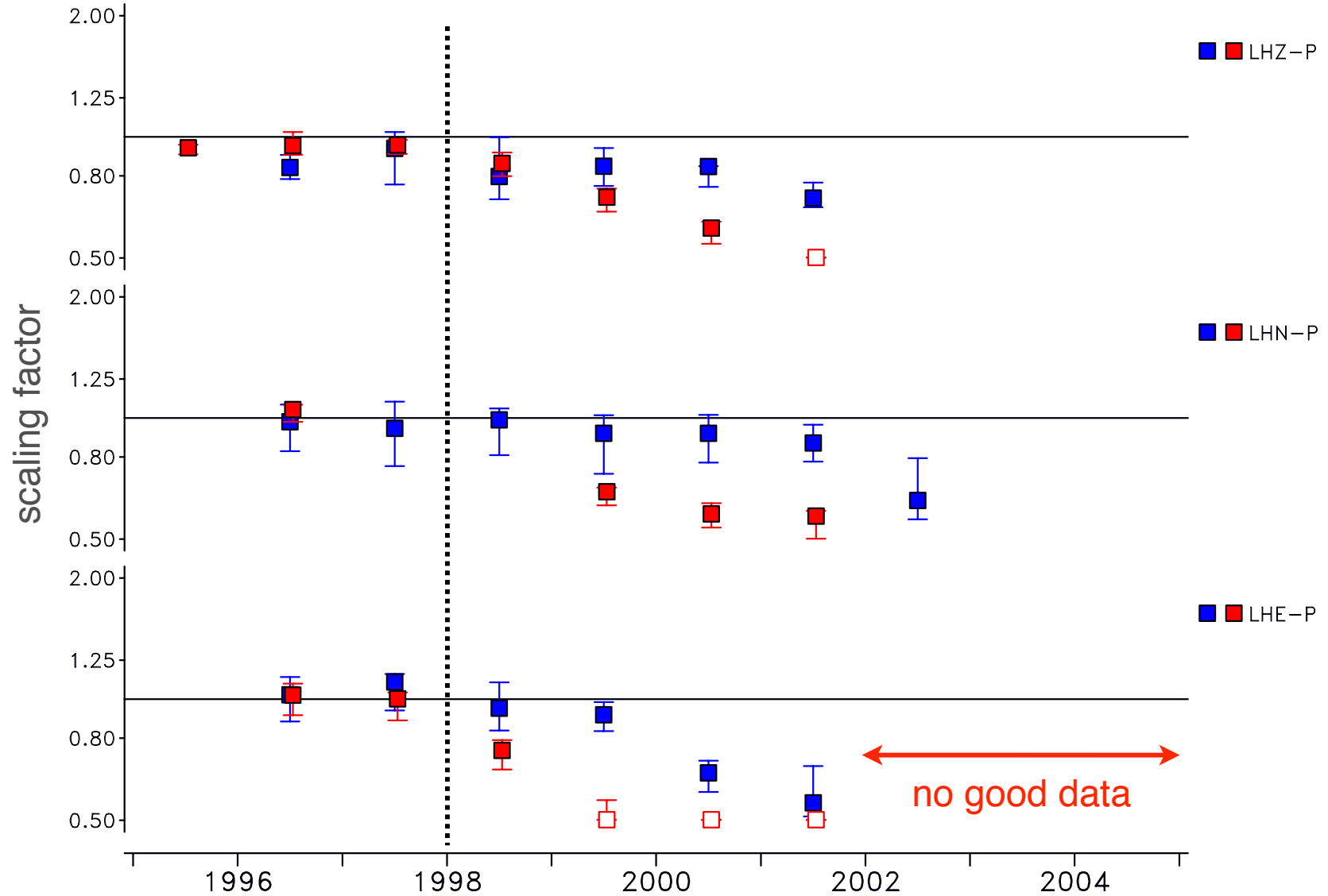
# Scaling factors at LVZ-II, 1993-2004

LVZ-II: Blue — Body; Red — Mantle



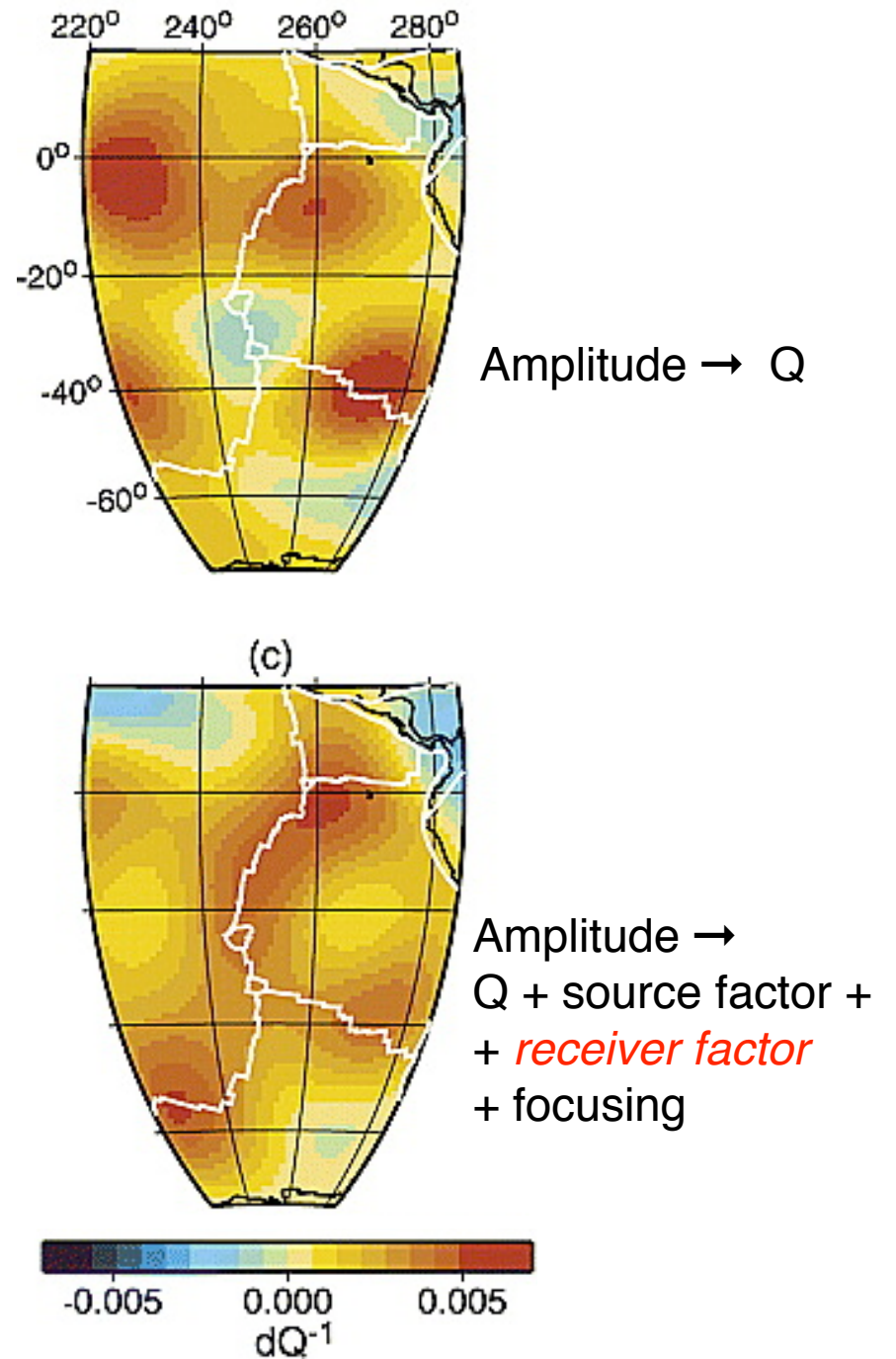
# Scaling factors at PEL-G, 1996-2002

PEL-G: Blue — Body; Red — Mantle



## Why does it matter?

- Amplitudes carry critical information for improving models of elastic and inelastic (Q) structure
- Also important for improvements in earthquake source modeling



(Dalton and Ekström, 2006)

*A simpler way to do this - if you have two instruments (A and B) in the same location:*

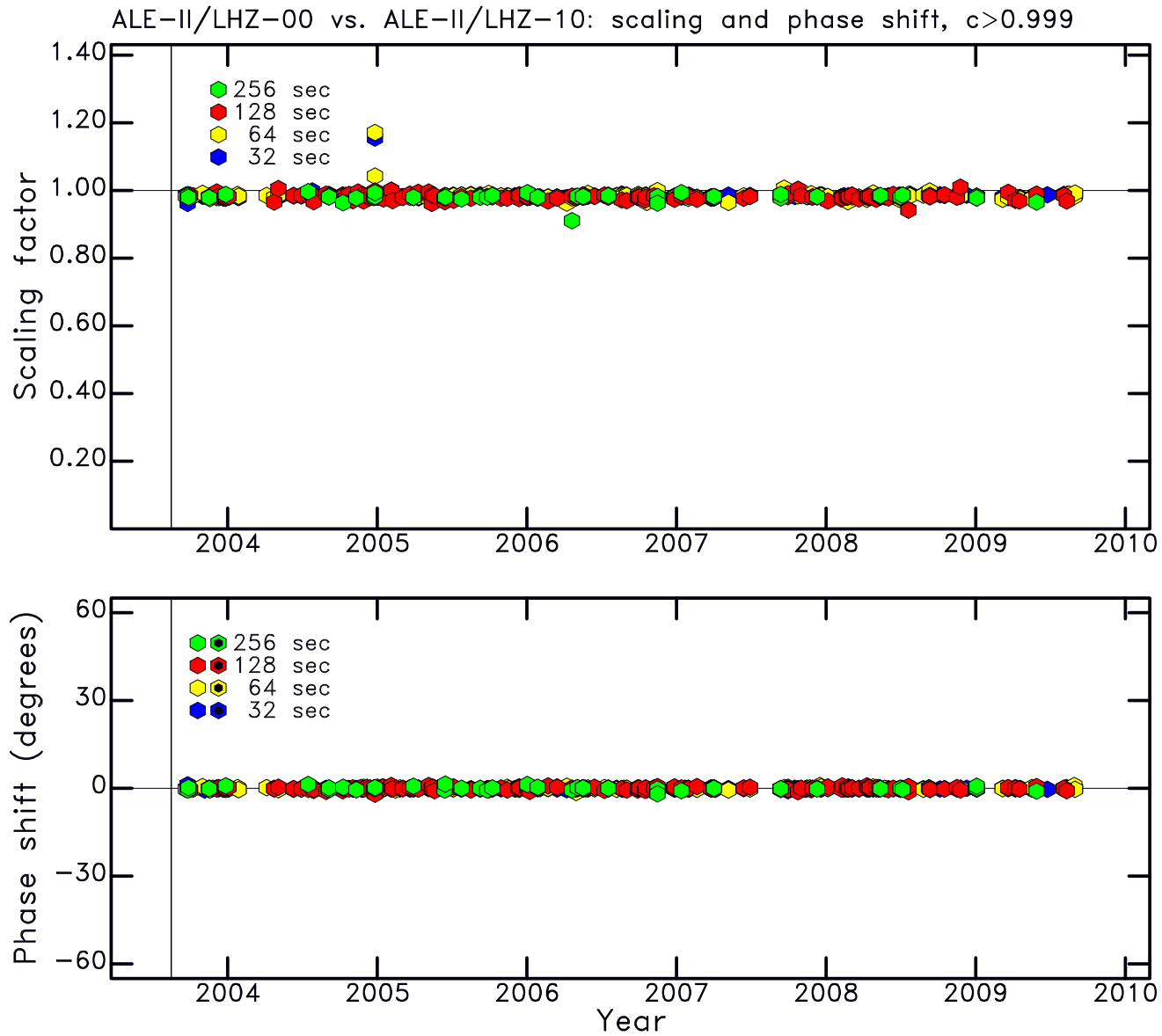
*calculate ratio of displacements at some period during times of high signal coherence*

$$\frac{\text{signal A}}{\text{response A}} = \text{displacement A} \quad (\text{deconvolution})$$

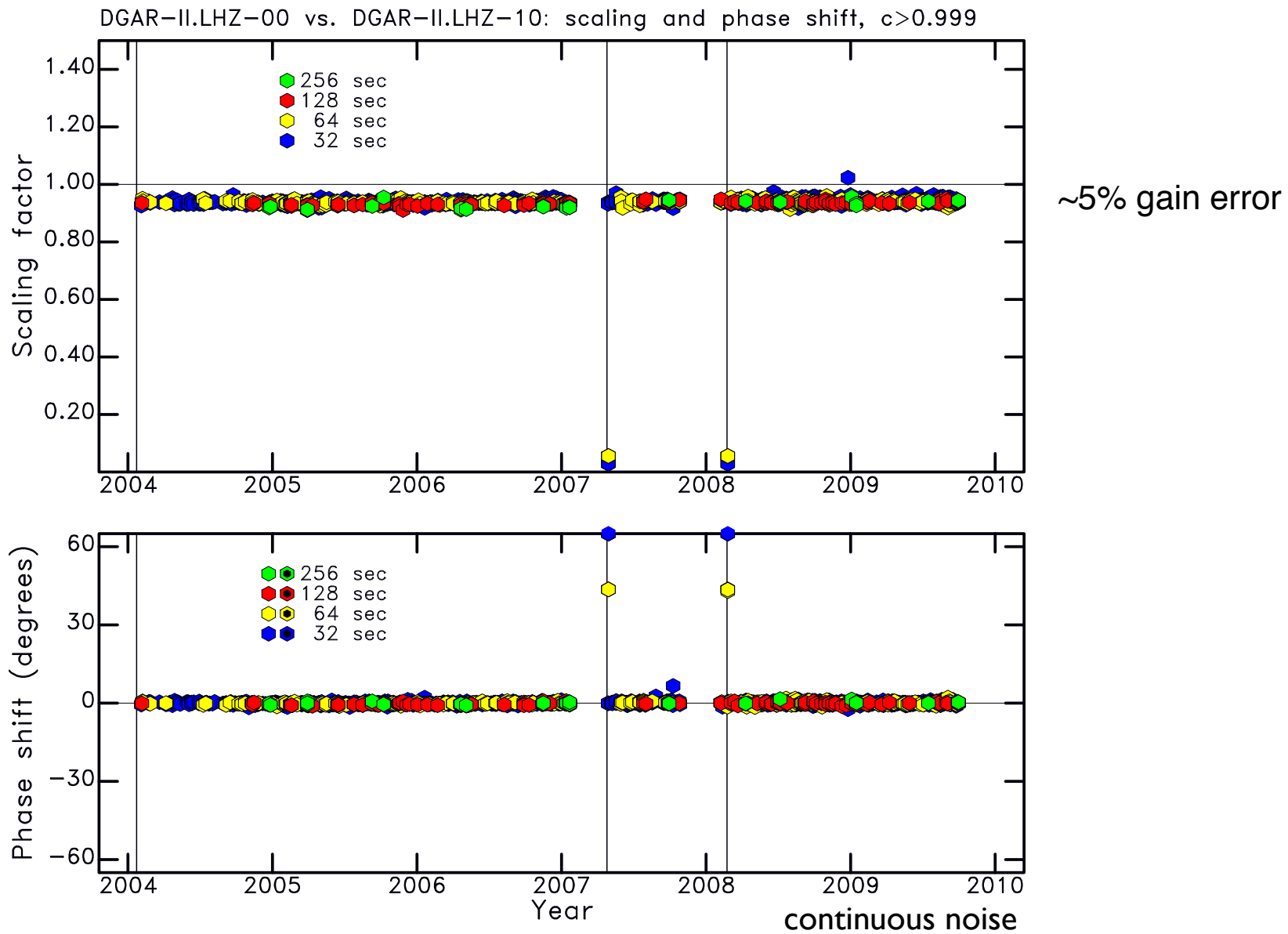
$$\frac{\text{signal B}}{\text{response B}} = \text{displacement B} \quad (\text{deconvolution})$$

$$\text{ratio} = \frac{\text{displacement A}}{\text{displacement B}} \quad \text{should be 1.0000!}$$

# Intersensor coherence, ALE-II LHZ, 2003-2009

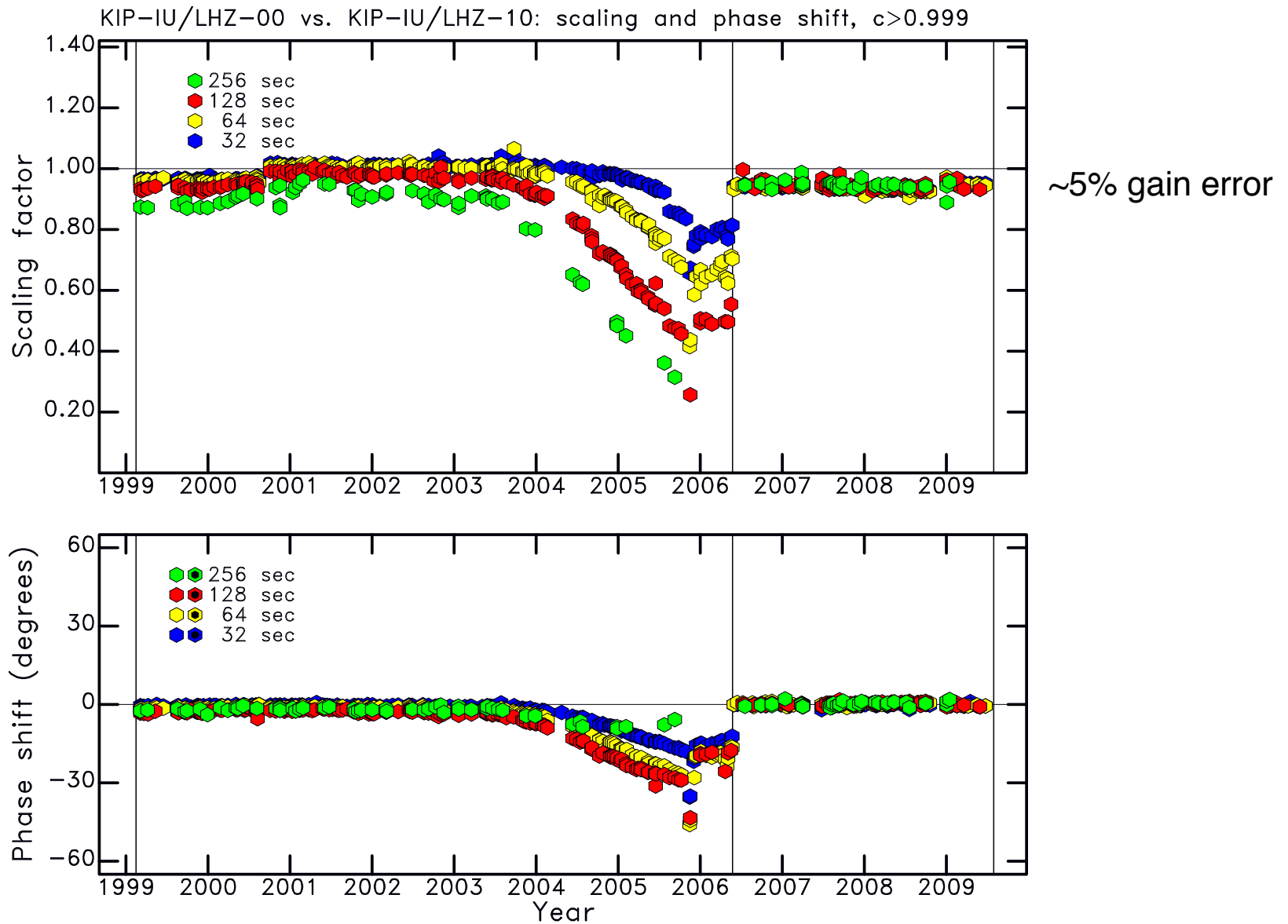


# Intersensor coherence, DGAR-II LHZ, 2003-2009



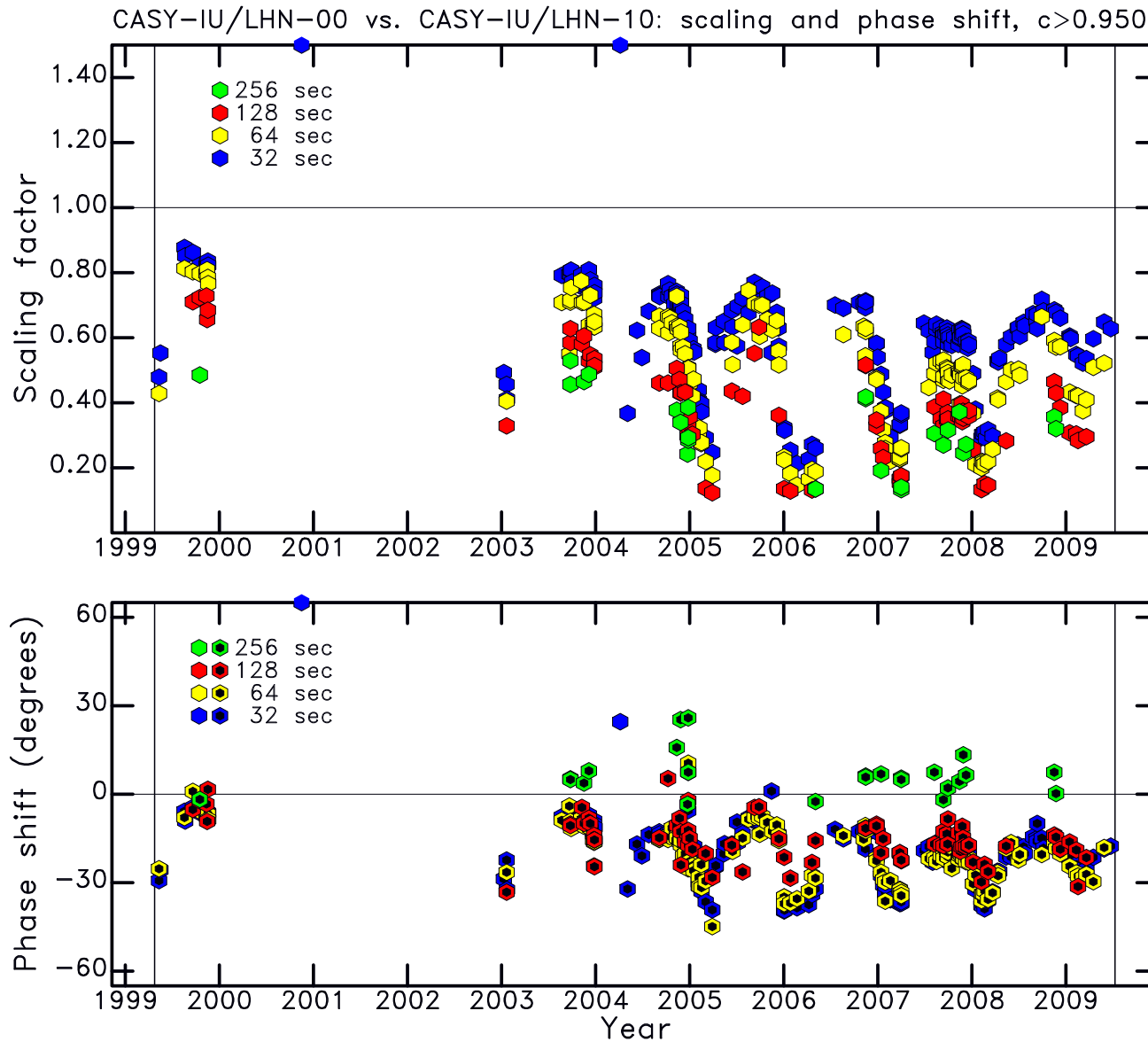


# Intersensor coherence, KIP-IU LHZ, 1999-2009



*STS-1 decay pattern*

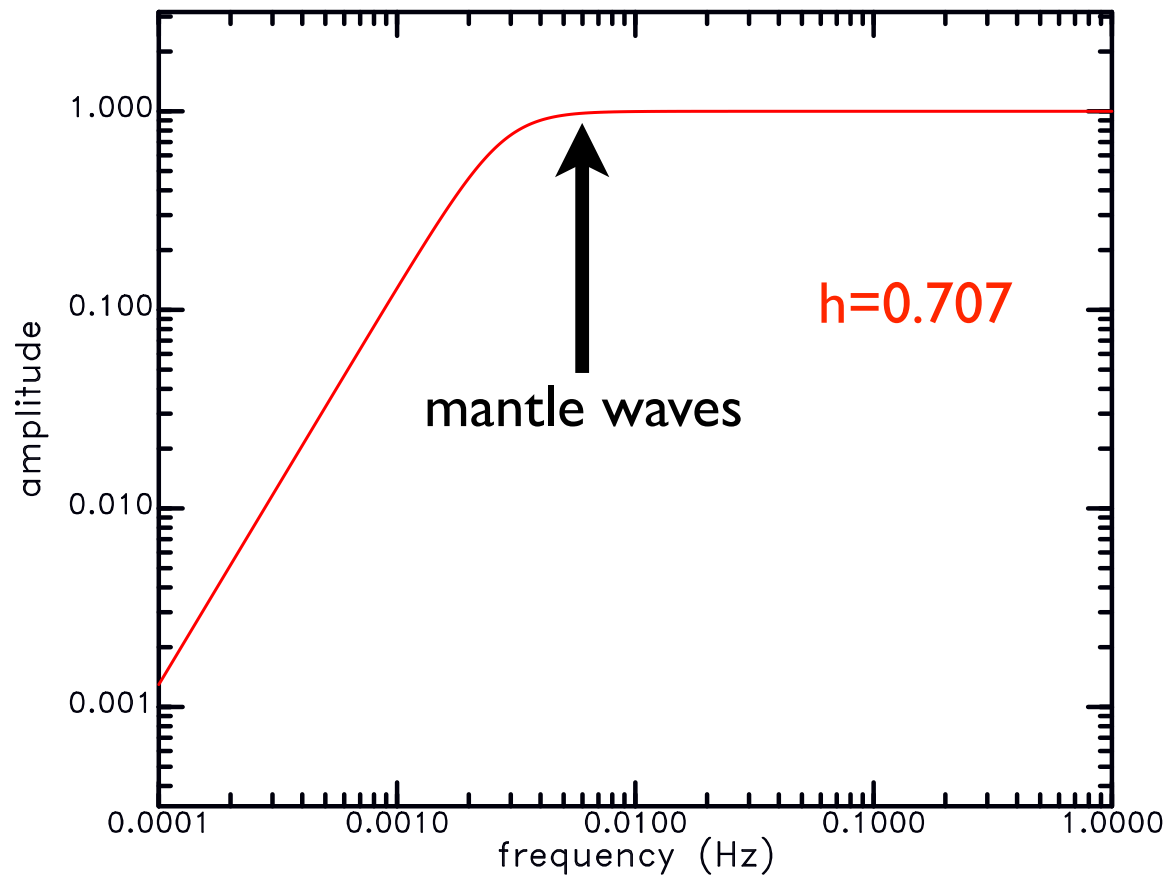
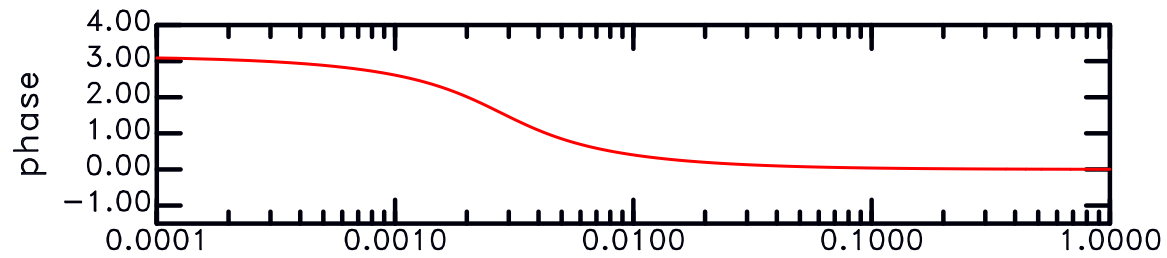
# Intersensor coherence, CASY-IU LHN, 1999-2009



*severe time- and frequency-dependent response error*

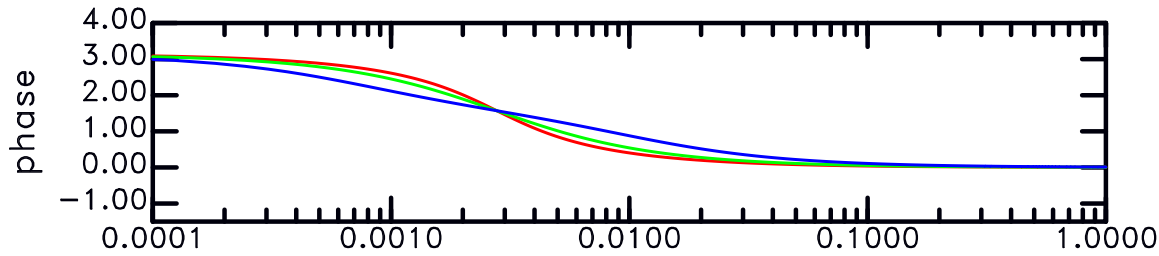
## STS-1 response decay

STS-1 generic response:  
360 second corner, critical damping ( $h=0.707$ )

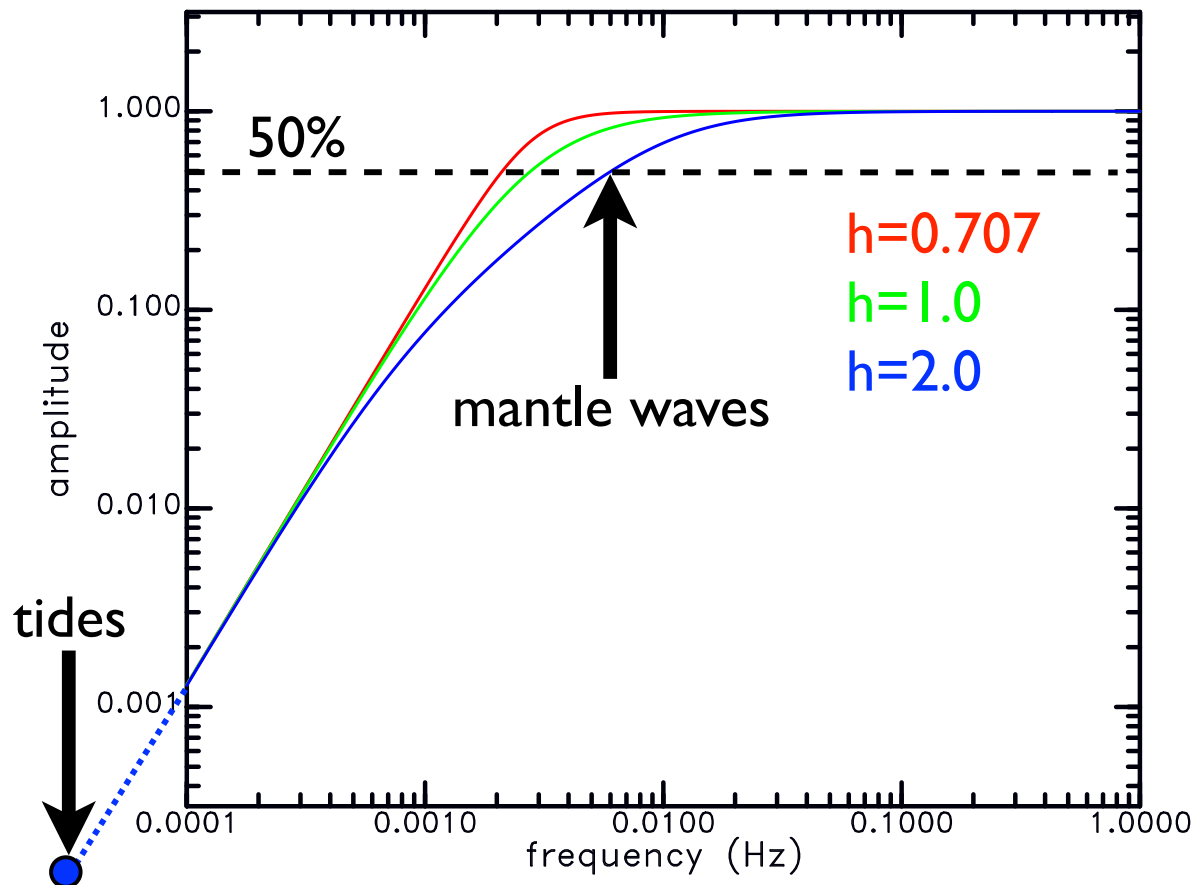


# STS-1 response decay

STS-1 typical corrupted response:  
360 second corner, overdamped



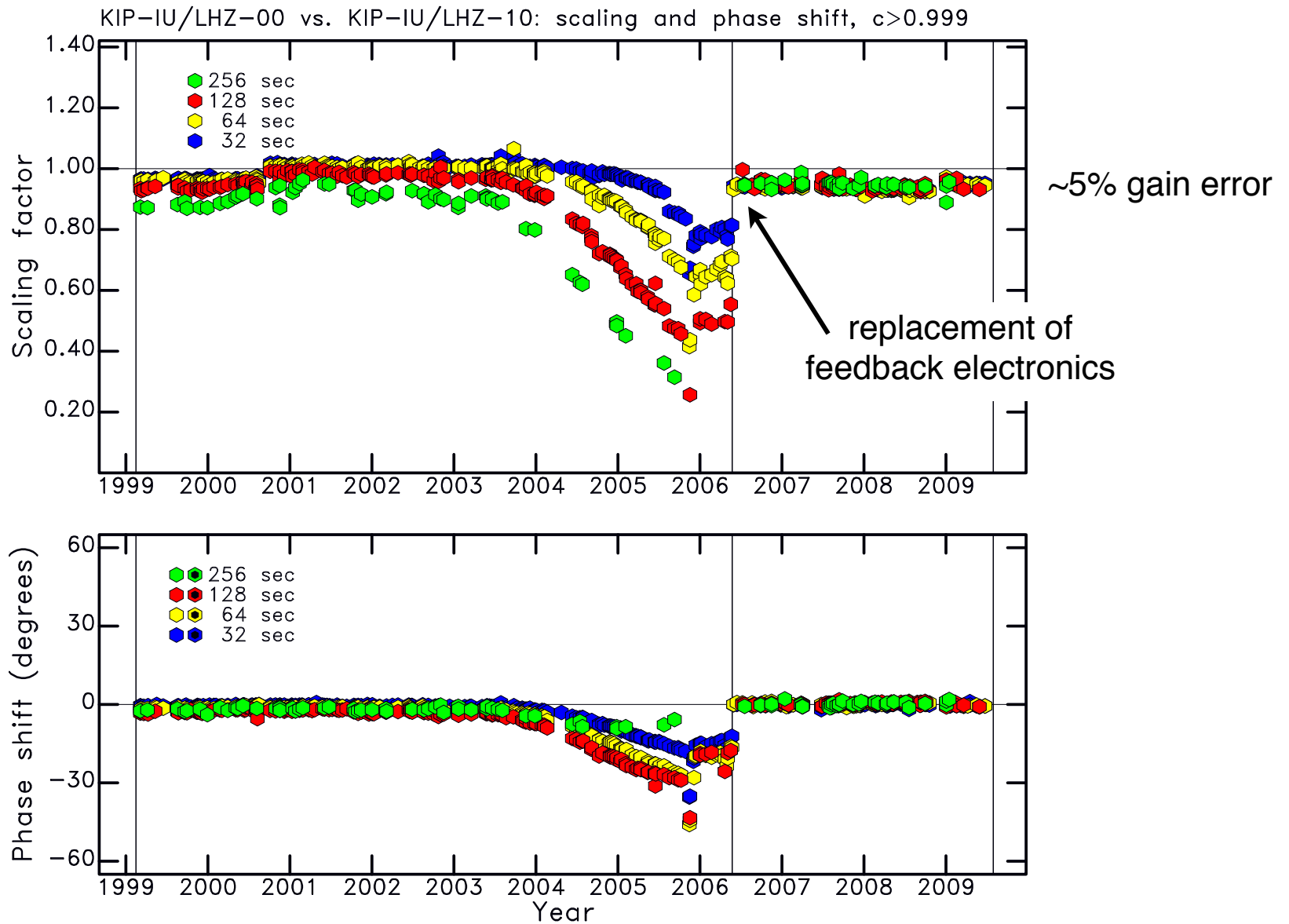
Hutt & Ringler:  
moisture in FBEs



Yuki & Ishihara:  
moisture in cable  
connectors

Hutt & Steim:  
too-short mechanical  
free period

# Intersensor coherence, KIP-IU LHZ, 1999-2009



*STS-1 decay pattern*

## Main points

1. The data can tell you a lot about your stations
2. Things change (calibrate!)
3. All networks can be improved

timing

orientation

response

noise level

All are important!